

World Space Observatory – Ultraviolet



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From Spanish ISSIS to Russian-Spanish-Mexican FCU



The Field Camera Unit (FCU) is one of the main scientific instruments of the WSO-UV mission. FCU will be the first UV-imager to be flown to a geosynchronous orbit (above the geocorona).

The main task of the FCU is to obtain high resolution images and low resolution spectra in Far- and Near- Ultraviolet bands.

The cooperation on FCU developments:

- INASAN and UCM scientific supervisors;
- IKI RAS prime contractor for FCU;
- Spain FUV detector supplier;
- Mexico optical elements (TBD).
- Japan coronagraph (TBD).

Russian-Spanish FCU



The preliminary design of FCU includes two channels.

1. FUV channel with MCP detector, **115-176 nm**:

- Solar blind detector;
- Diffraction-limited imaging in FUV;
- High sensitivity in photon-counting mode;
- High time resolution;
- Low resolution field spectroscopy.
- 2. NUV channel with CCD detector, **174-310mn**:
 - Wide field of view;
 - High dynamic range;
 - High angular resolution;
 - Low resolution field spectroscopy;
 - Possibility for extended spectral range: 115-1000 nm.

Russian-Spanish FCU



- Now IKI RAS (Russia) and NAOJ (Japan) are looking for a possibility to realize the third channel with a coronagraph mode.
- More details about the coronagraph mode will be presented by Alexander Tavrov (IKI).





Russian-Spanish FCU



Main features of FCU

Parameters	ISSIS	Far-UV	Near-UV
Detector	MCP, analogue of UVIT (Spain)	MCP type detector	CCD
Spectral range, nm	115-310	115-176	174-310 (115-1000)
Effective area, M²	0.054	0.068	0.27
Field of view, arcsec×arcsec	70×75	121×121	597×451
Angular resolution, arcsec	0.03	0.08	0.146
Detector size, mm	40	30	49×37
Number of filters	2 x (5+2 neutral)	Up to 10	Up to 15