

ULTRAVIOLET ASTRONOMY IN THE XXI CENTURY



e-Workshop 2020 – October 27-29

Clues on the afterlife galaxies from SDSS spectra and GALEX photometry

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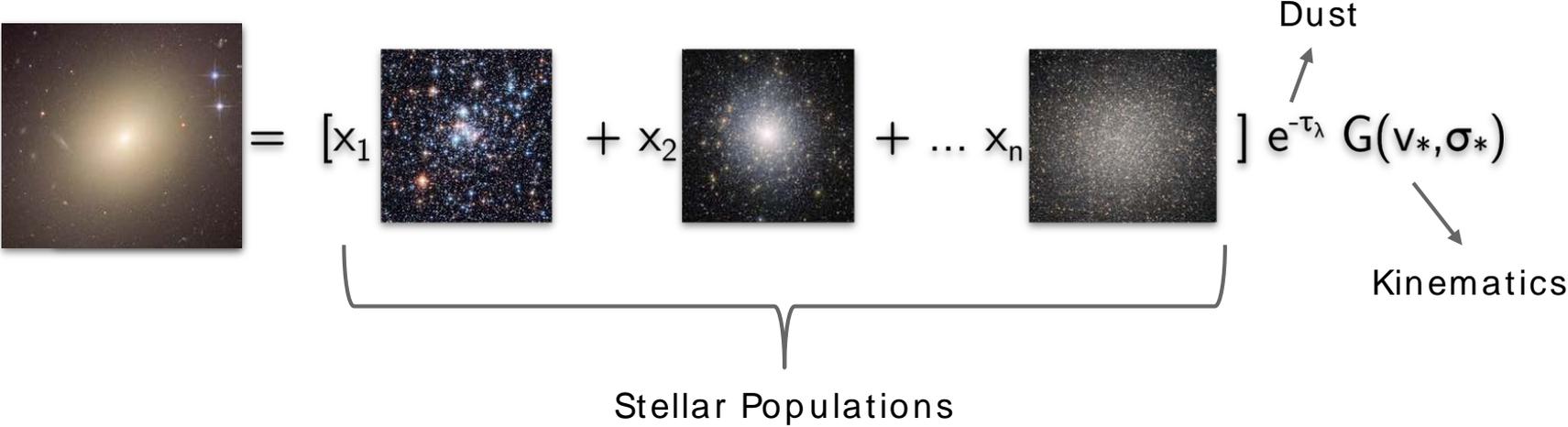


OSSERVATORIO
ASTRONOMICO DI PADOVA

A background image of a starry night sky. The sky is dark with numerous stars of varying brightness. A prominent feature is a nebula or galaxy structure, showing a mix of blue and brownish-orange colors, with a bright, glowing region on the left side. The text 'Spectral synthesis: STARLIGHT' is overlaid in white, sans-serif font, centered horizontally and vertically.

Spectral synthesis:
STARLIGHT

Spectral synthesis

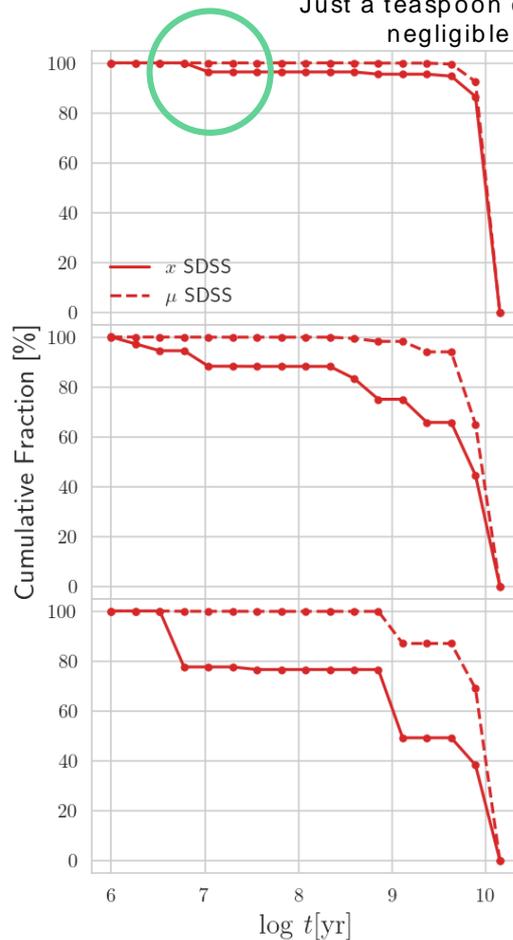
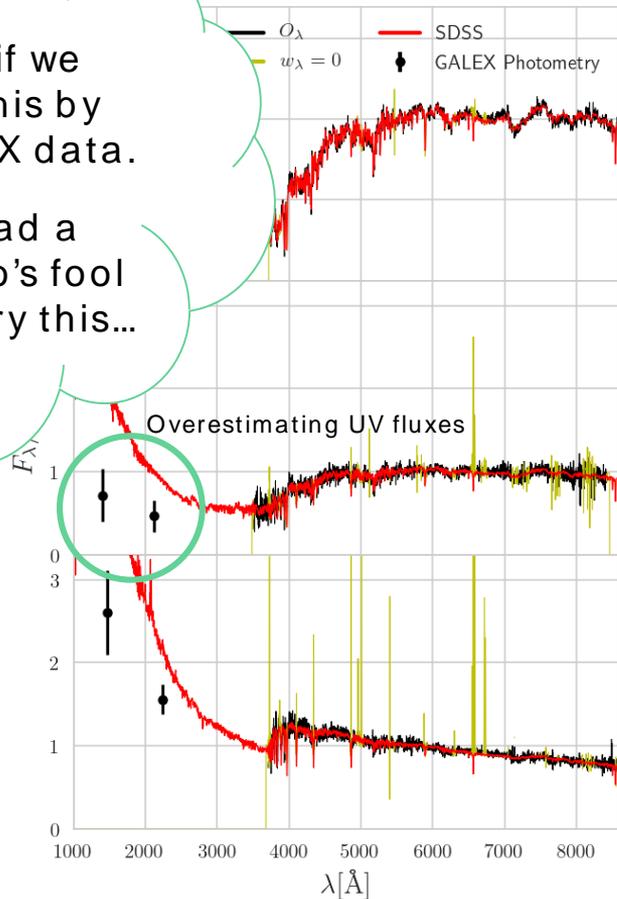


An old known issue

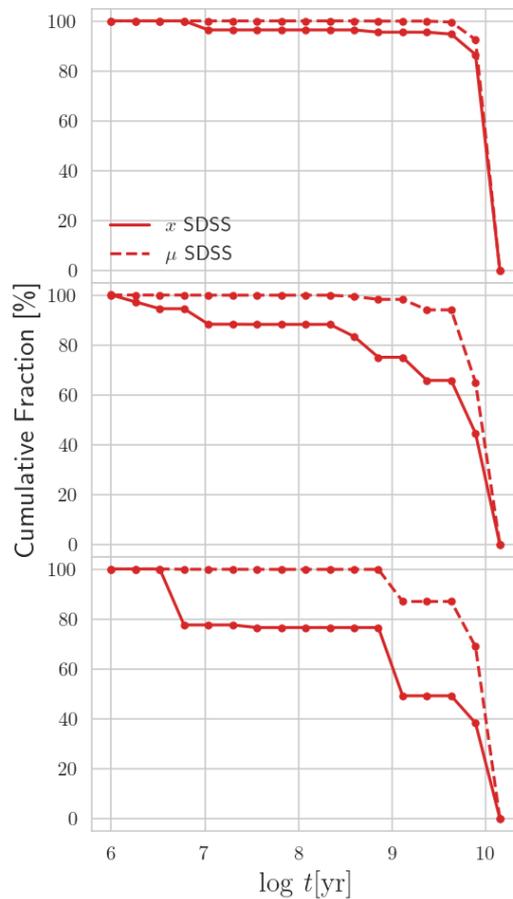
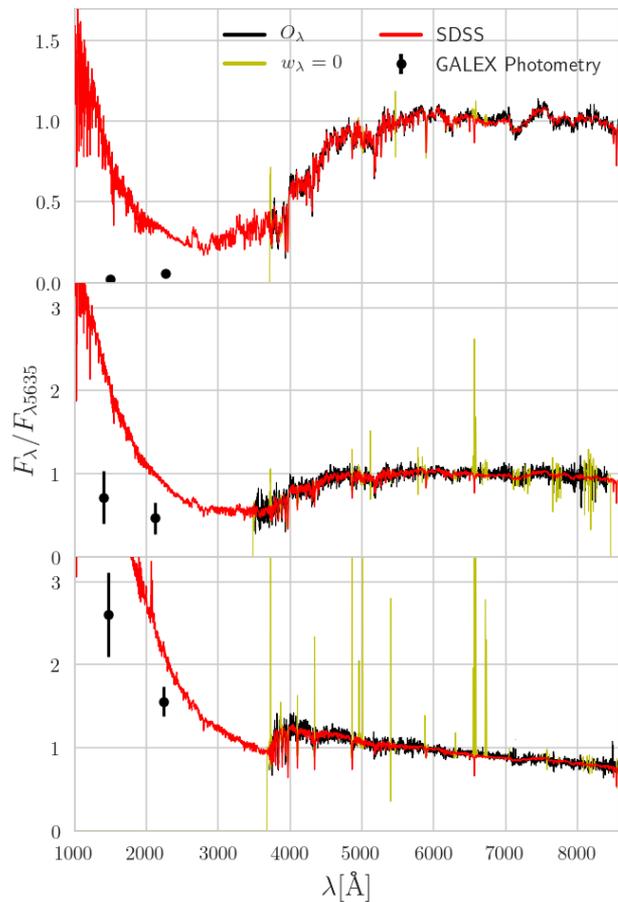
Just a teaspoon of optical light,
negligible in mass

I wonder if we
could fix this by
fitting GALEX data.

If only I had a
student who's fool
enough to try this...

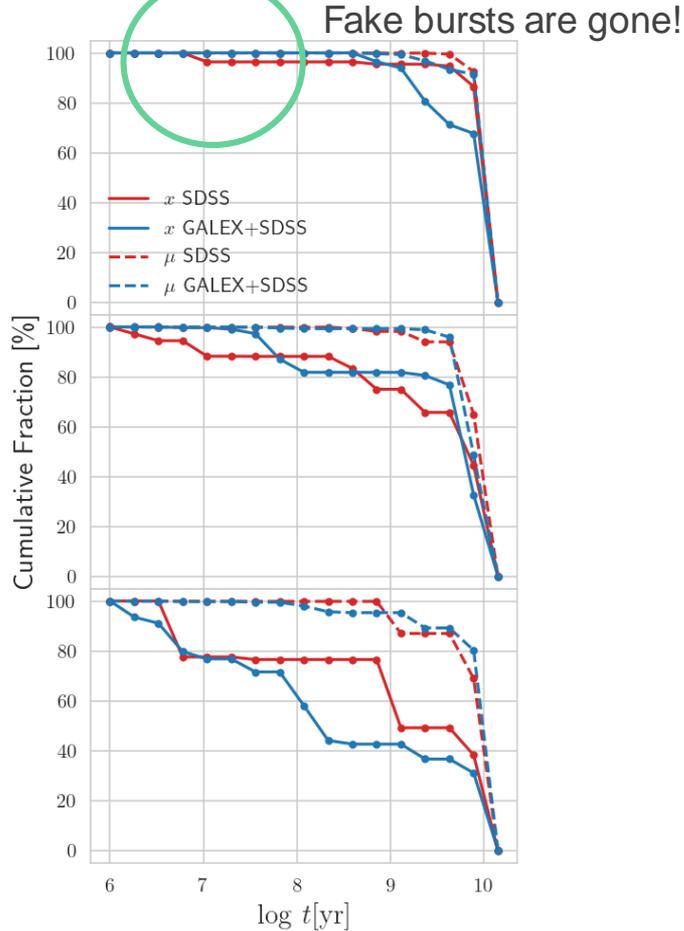
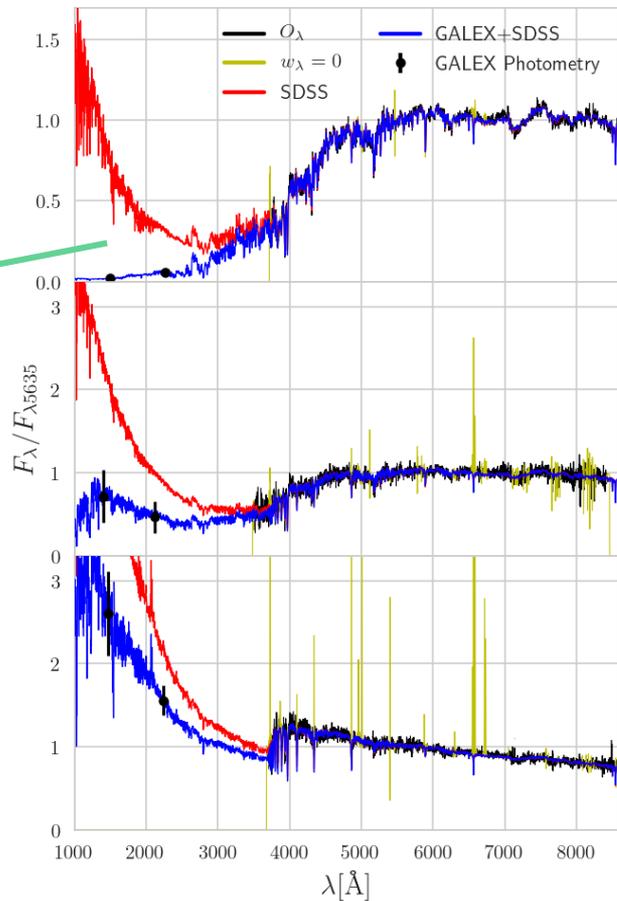


GALEX+SDSS synthesis



GALEX+SDSS synthesis

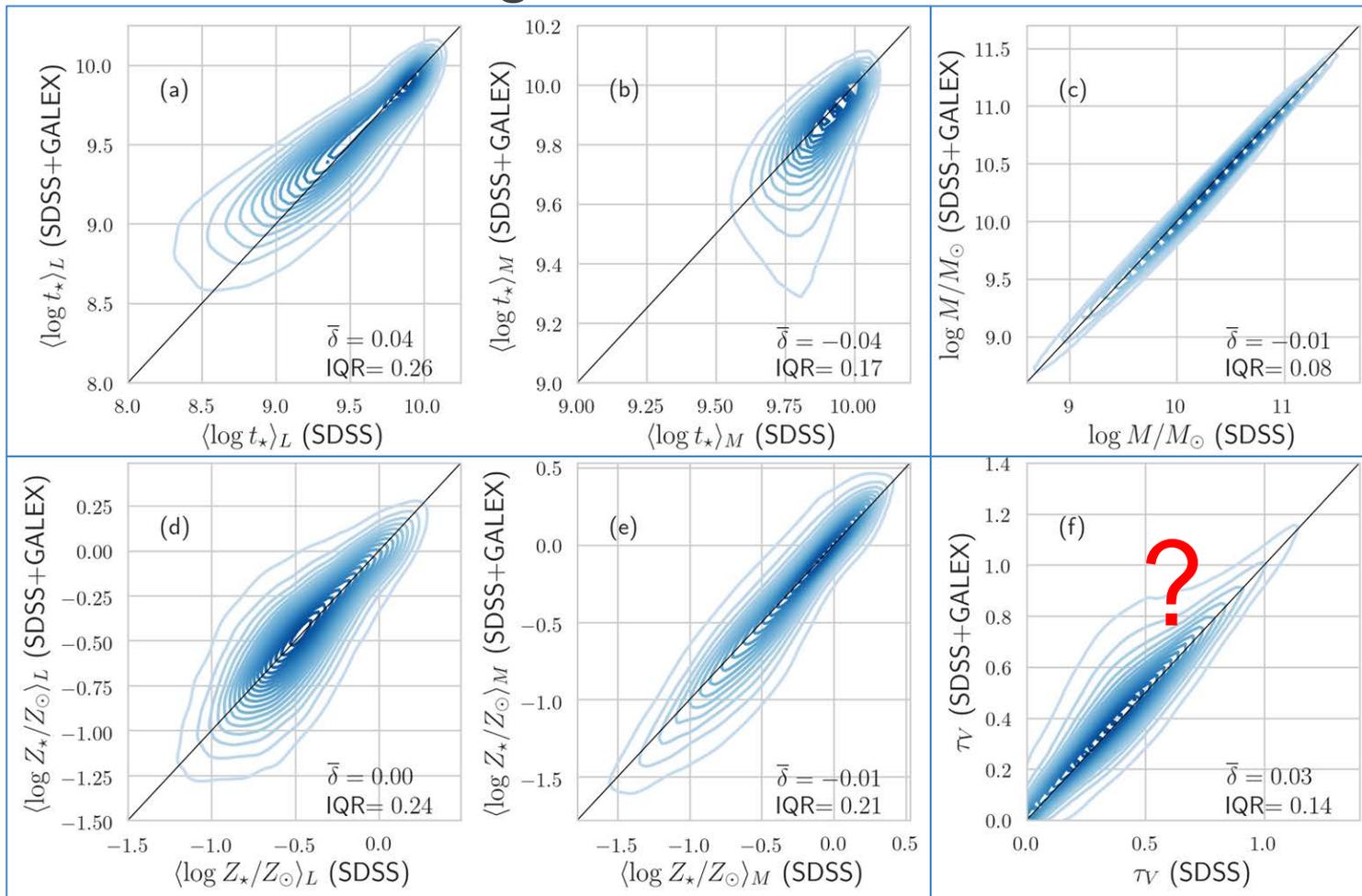
GALEX data prevents the overshooting of UV fluxes!





What changes
with UV data?

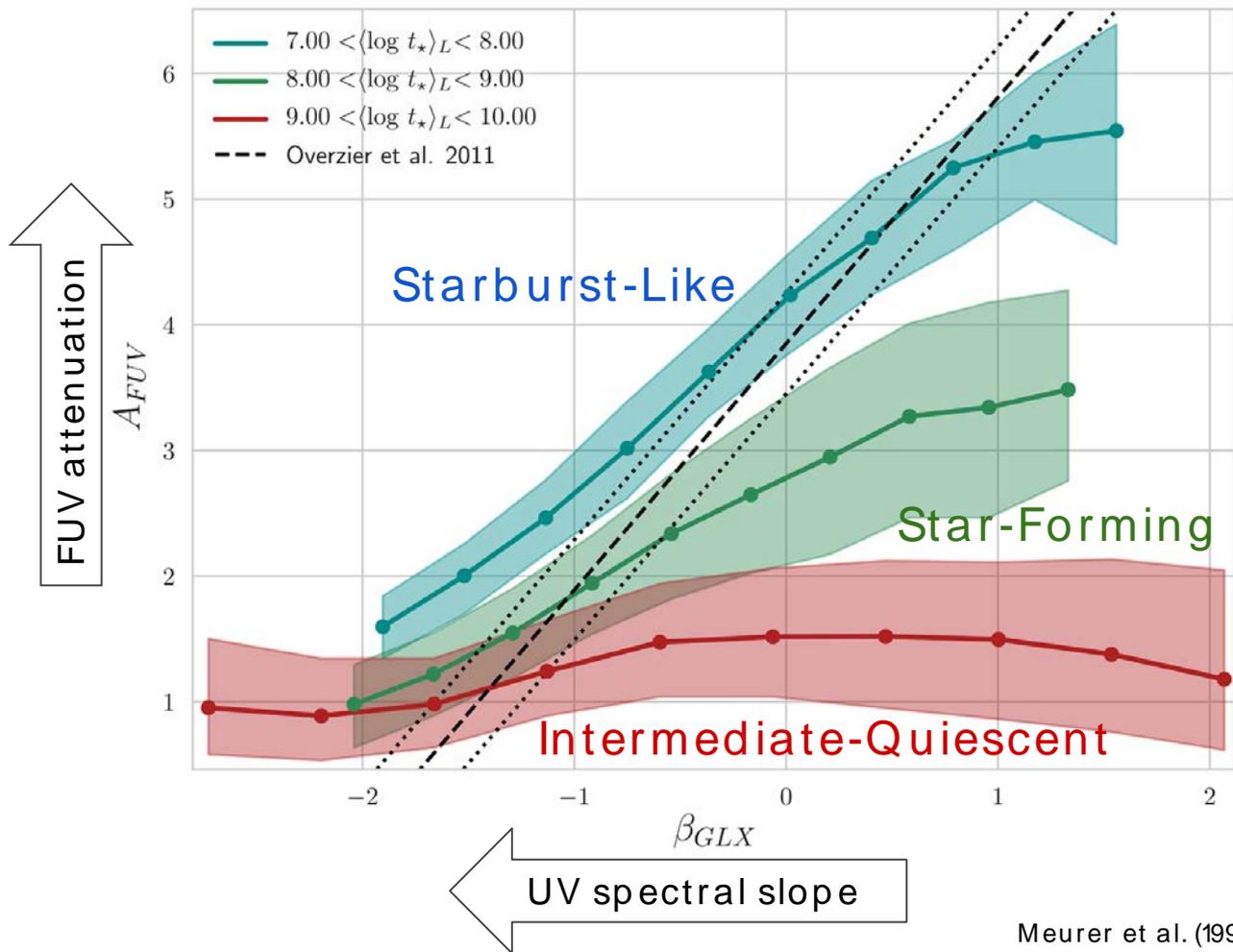
What changes with GALEX data?



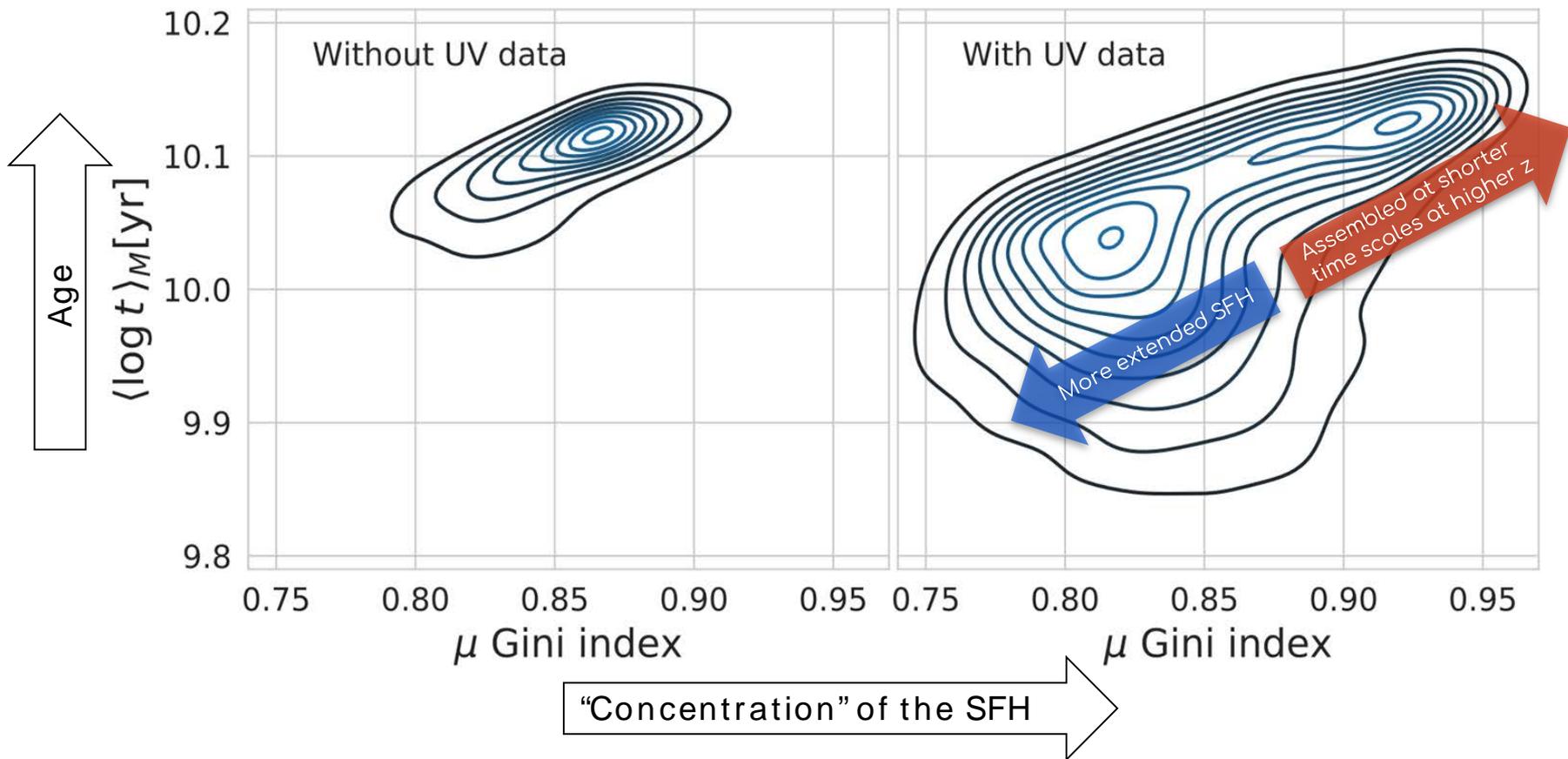
$\beta - A_{FUV}$ relation

β is the UV power-law slope obtained by fitting the UV spectrum with:

$$F_{\lambda} = \lambda^{\beta}$$



A bimodality in the red sequence





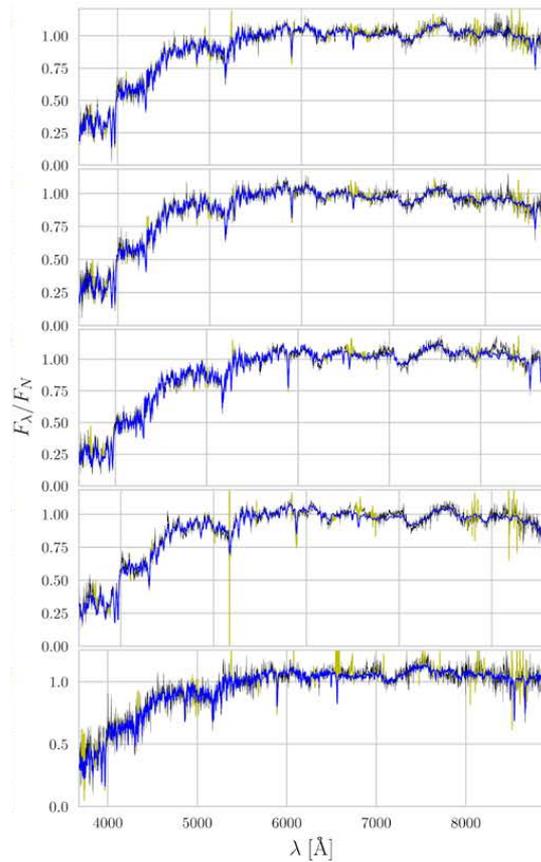
Unveiling the
afterlife of
galaxies

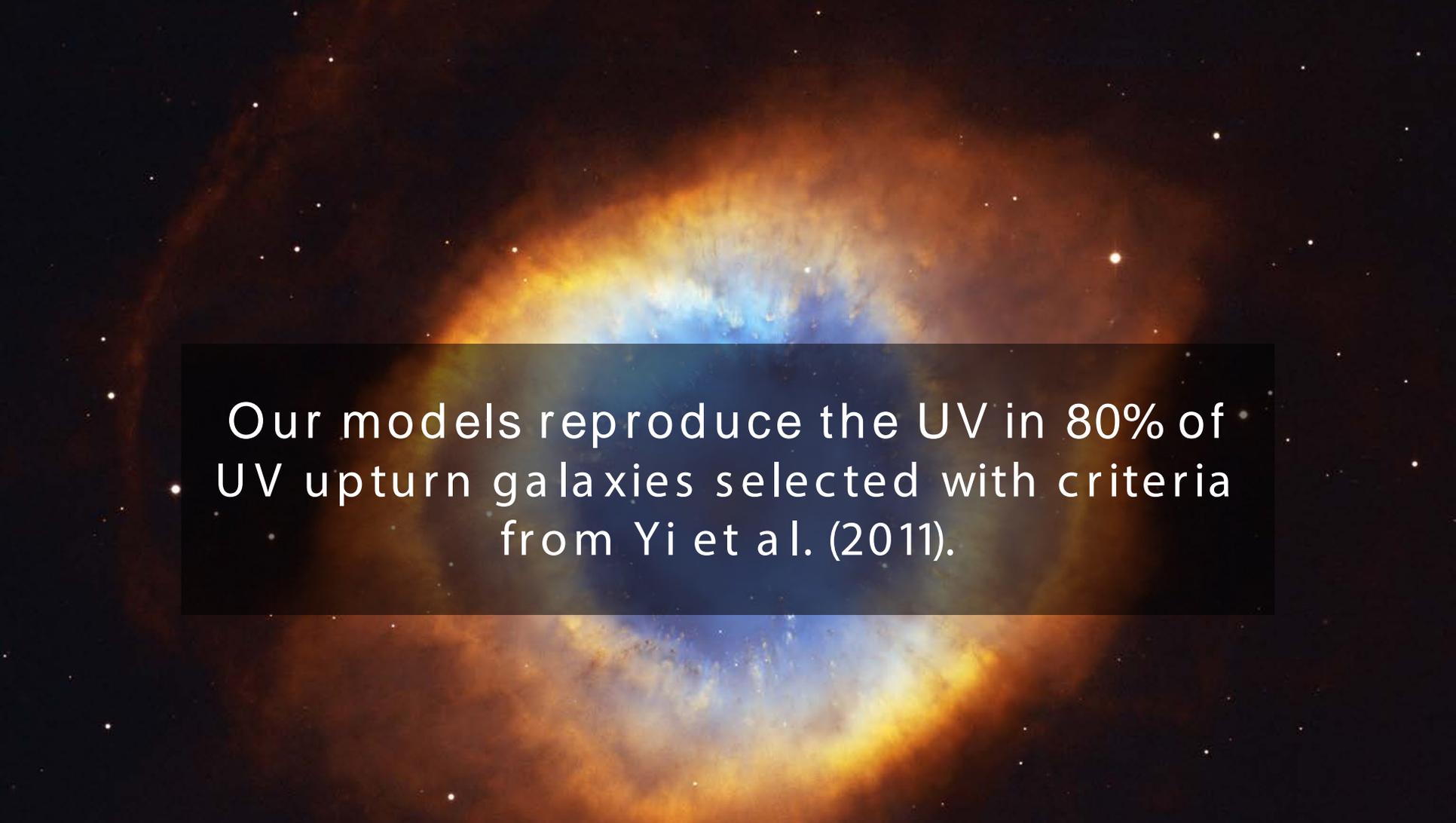
What will you be looking at?

3453 galaxies:

- With SDSS spectra
- Detected in both GALEX bands (NUV and FUV)
- Classified as elliptical by galaxy zoo
- In the **red sequence** ($NUV-r > 5$)

Spectral synthesis of ETGs



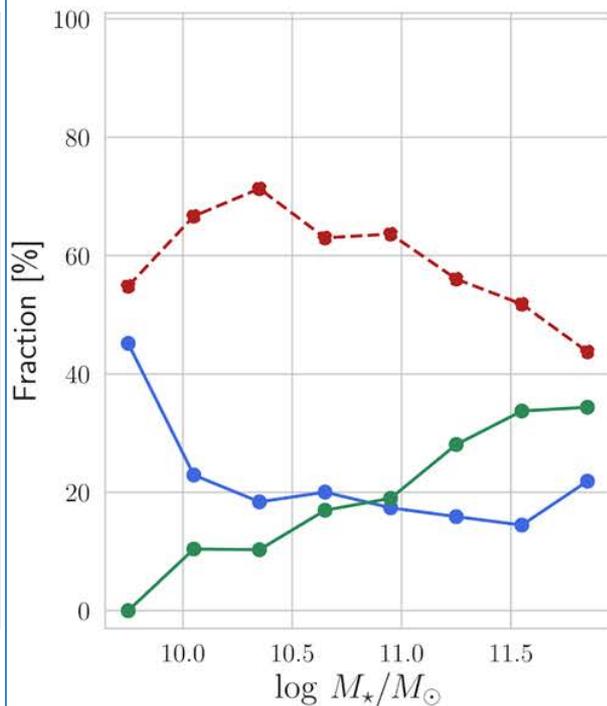
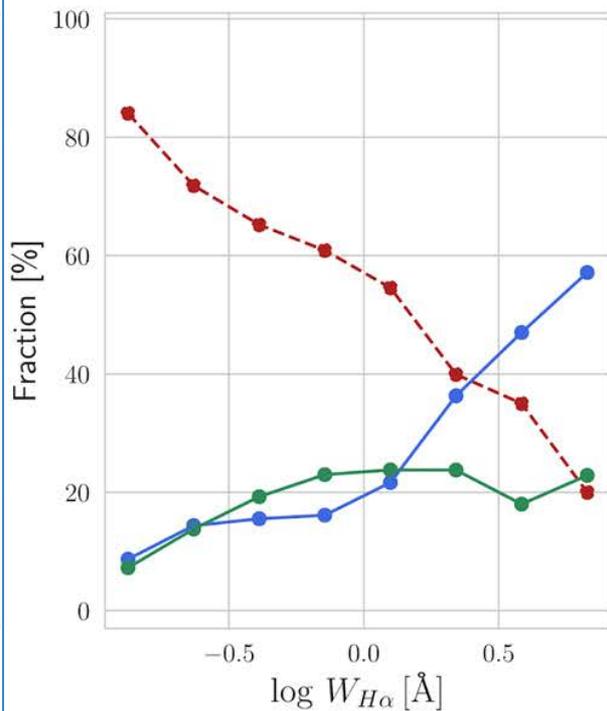
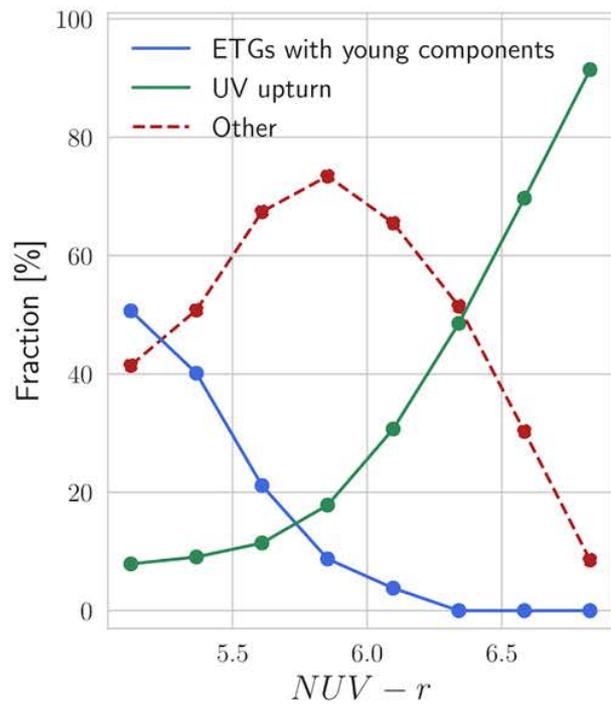


Our models reproduce the UV in 80% of
UV upturn galaxies selected with criteria
from Yi et al. (2011).

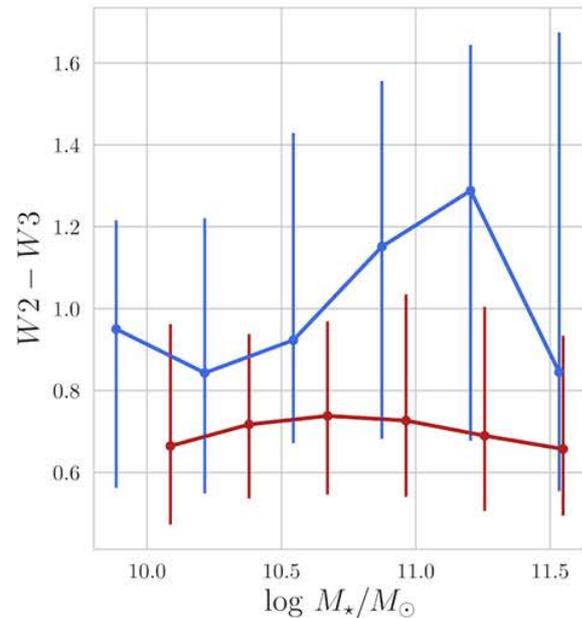
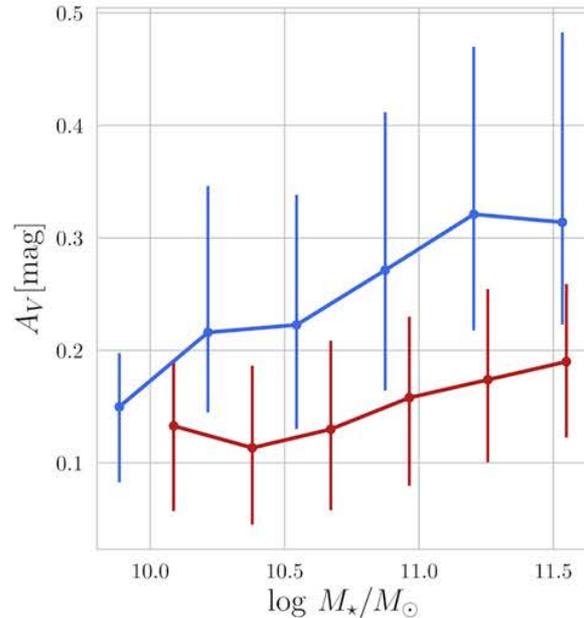
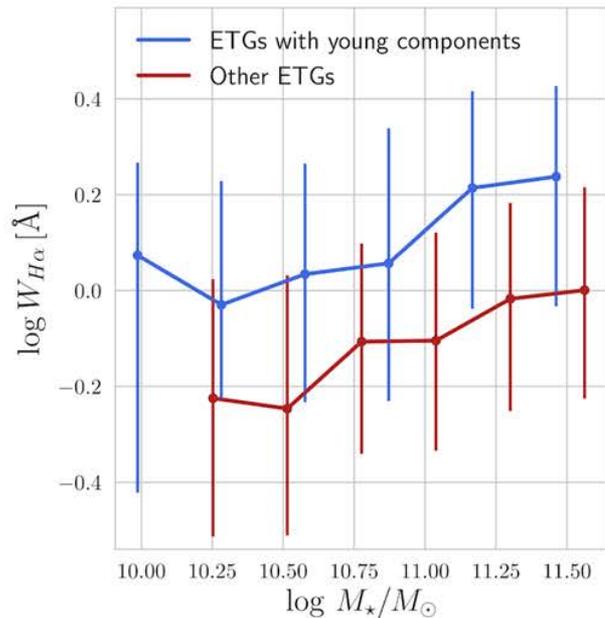
Sub-classes of ETGs

- ETGs with young components: At least 5% of FUV emission from stars younger than 1 Gyr (17.6%)
- UV upturn: No young components and $FUV-NUV < 0.9$ (21.6%)
- Other: None of the above (60.6%)

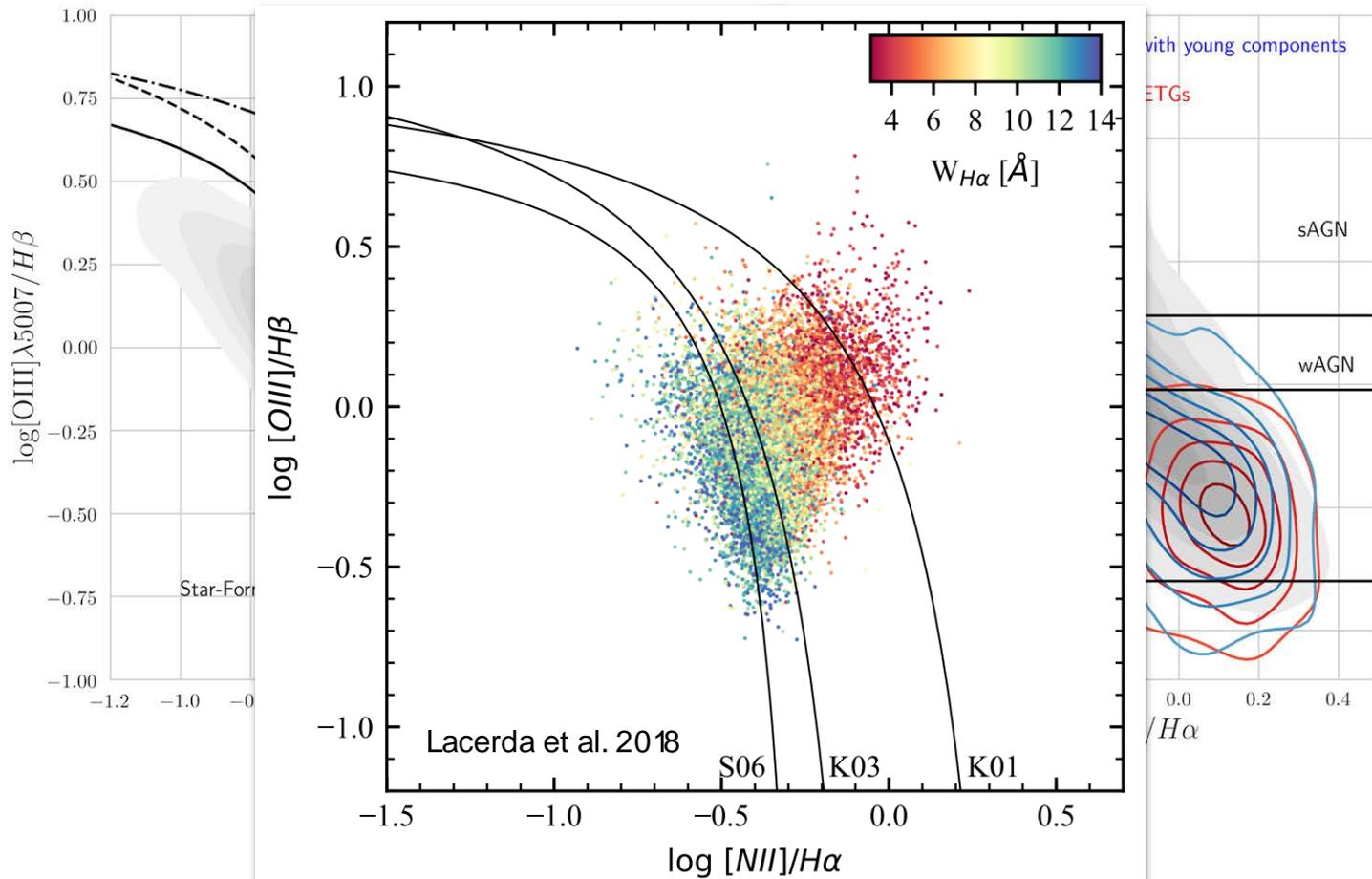
Distributions of the sub-classes



The ISM of ETGs with young components



Ionization sources: HOLMES or SF?

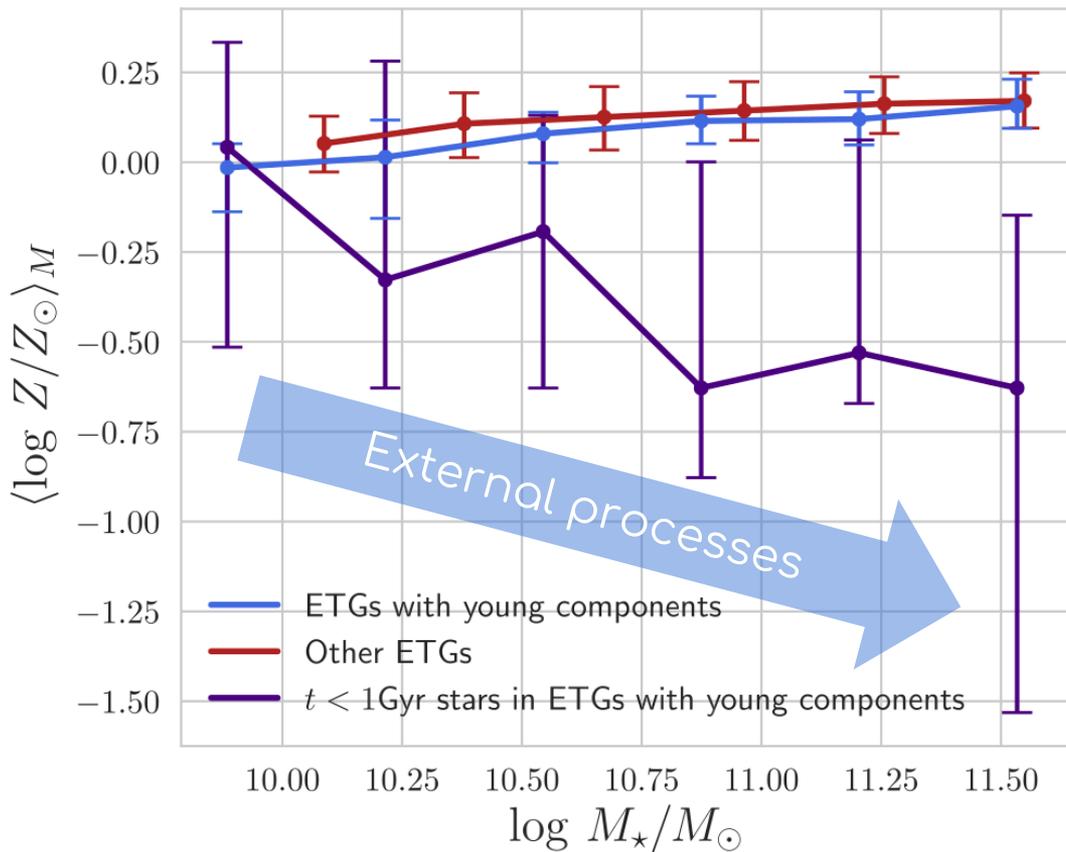




Beware! Convoluted
plot ahead.

How do young components form?

Downsizing
Internal processes
Moving up towards
the **red sequence**



Rejuvenation
External processes
Moving towards or
away from the **red
sequence**



Thank you!