

ULTRAVIOLET ASTRONOMY IN THE XXI CENTURY

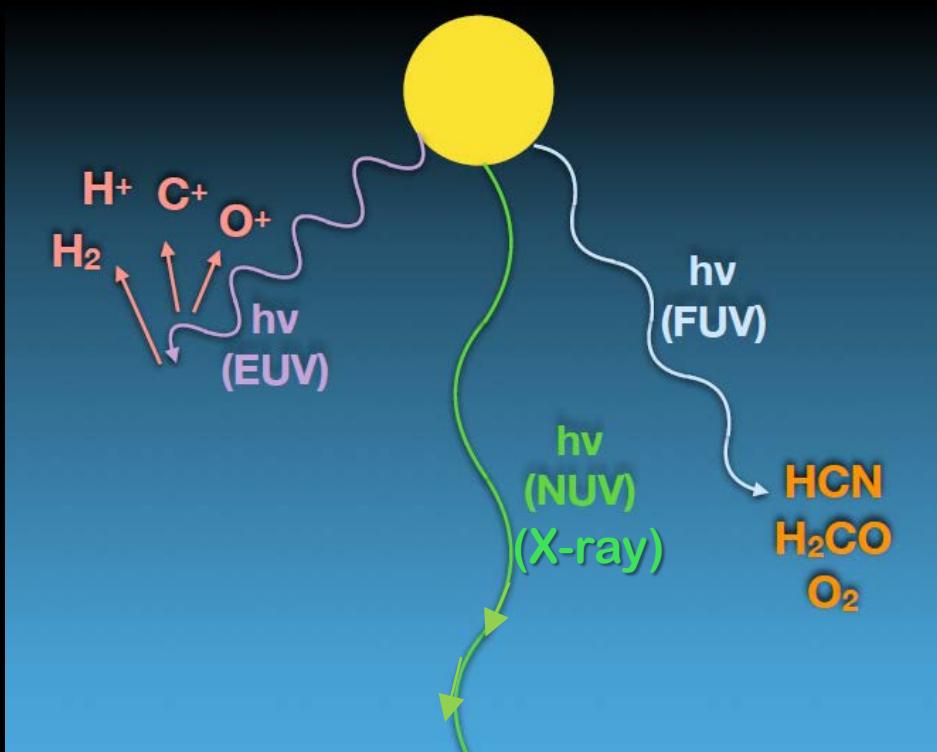
e-Workshop 2020 – October 27-29

Atmospheric Escape from Extrasolar Planets: From stellar inputs to exoplanetary signatures with the ESCAPE and CUTE missions



Kevin France
University of Colorado
NUVA 2020

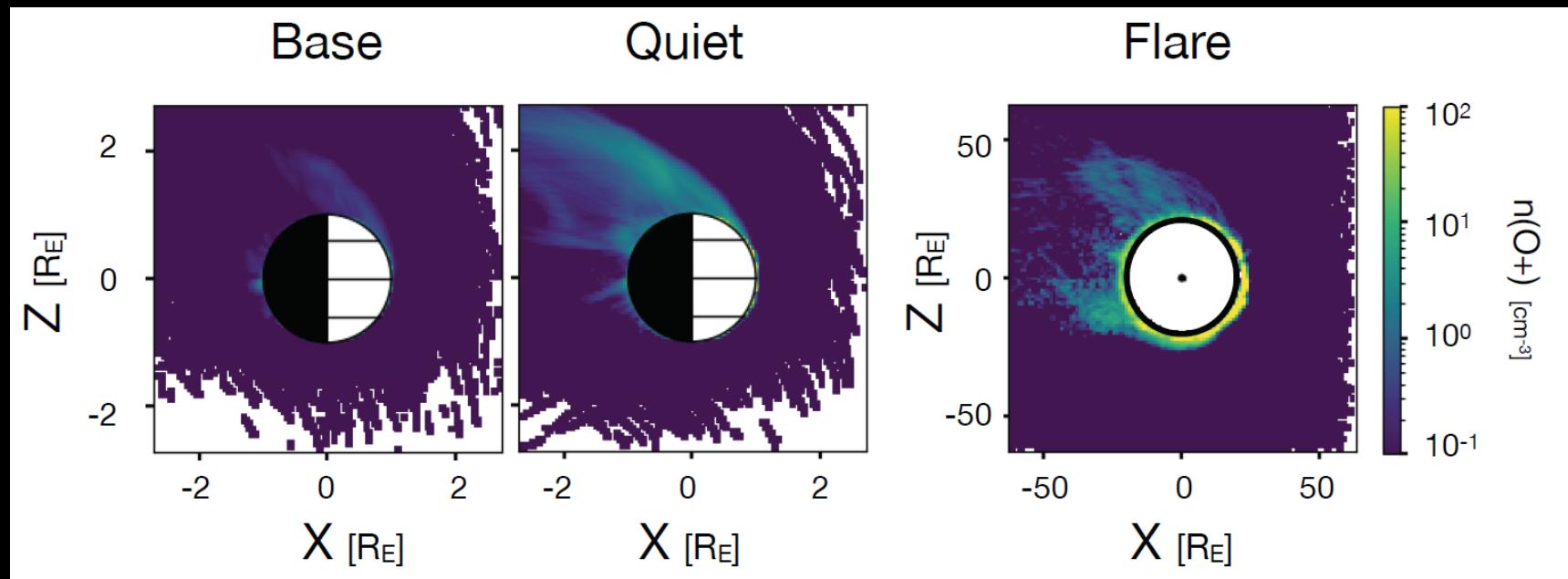
Introduction: Star-Planet Interactions and Atmospheric Impacts



- Photons of different energy play distinct roles, and all contribute to the observable signatures of that atmosphere
- The high-energy stellar emission dominates atmospheric photochemistry, ionization, and heating

Figure courtesy of
Paul Rimmer - Cambridge

Introduction: Star-Planet Interactions and Atmospheric Impacts

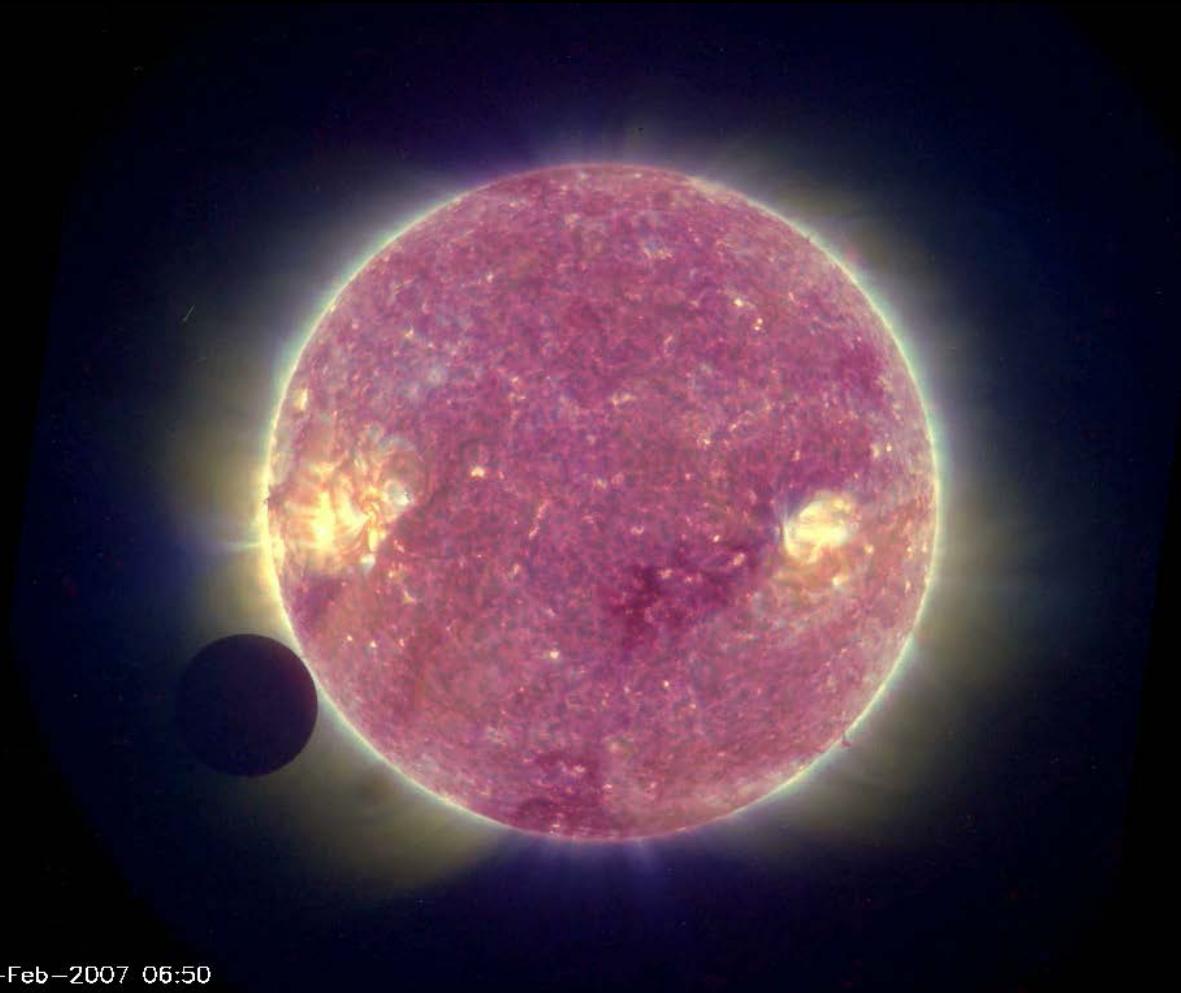


- Example: Rapid atmospheric escape driven by a combination of stellar EUV photons and particles (stellar winds, CMEs) may prevent exoplanets from maintaining habitable environments over geologic timescales



CREDIT: Efrain Morales Rivera, Astronomical Society of the Caribbean, PHL @ UPR Arecibo

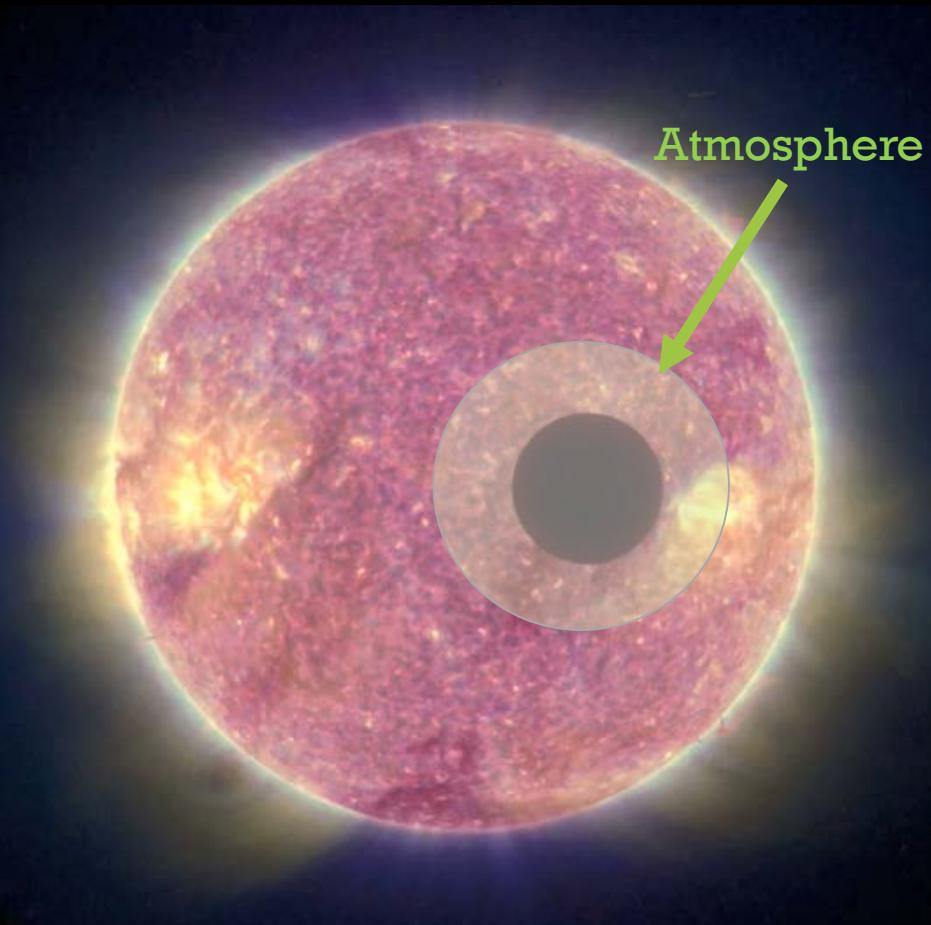
STUDYING ATMOSPHERIC ESCAPE TODAY



Occultation
Depth =
 $(R_p / R_*)^2$

EXOPLANET ATMOSPHERES

- Narrow-band/spectroscopic transit analysis can probe absorption by specific atmospheric constituents



Occultation
Depth =
 $(R_p(\lambda) / R_*)^2$

Transit Spectroscopy:
in-transit vs. out-of-transit

- Composition
- Temperature structure
- Velocity flows
- Mass-loss rates



Laboratory for Atmospheric and Space Physics
University of Colorado Boulder



ARIZONA



Trinity College Dublin



The University of Dublin



BUTE



NASA



Université
de Toulouse



ÖAW
Austrian Academy
of Sciences



The Colorado Ultraviolet Transit Experiment (CUTE)

University of Colorado:

Kevin France (PI), Brian Fleming (PS), Arika Egan*, Rick Kohnert (PM), Nicholas Nell*, Stefan Ulrich, Nick DeCicco, Ambily Suresh

United States:

Tommi Koskinen (UofA), Matthew Beasley (SwRI), Keri Hoadley (Caltech/Iowa)

Europe:

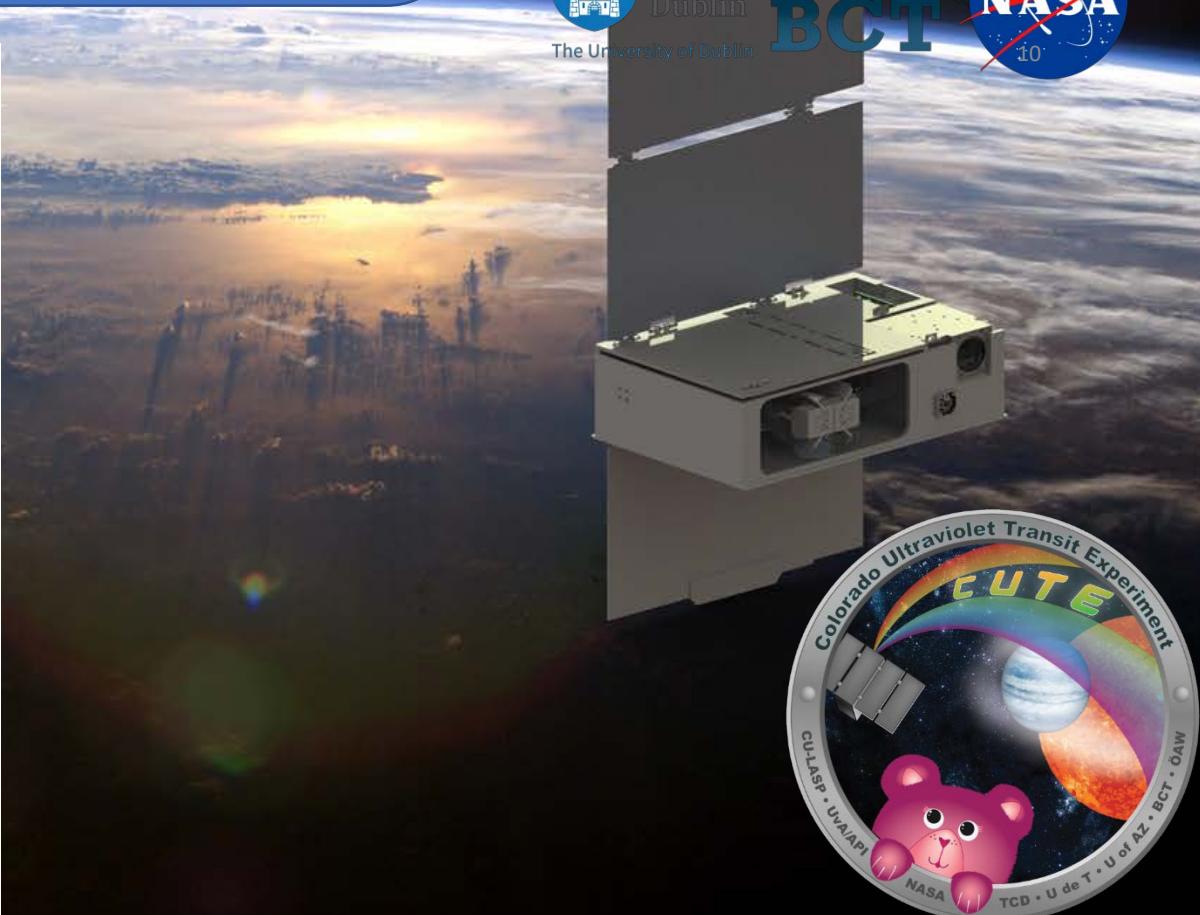
Jean-Michel Desert (Amsterdam), Luca Fossati (ÖAW), Pascal Petit (UdeT), Aline Vidotto (TCD)

CUTE:

11.0 cm x 23.7cm x 36.2 cm

Family Size Cheerios available
on Walmart.com:

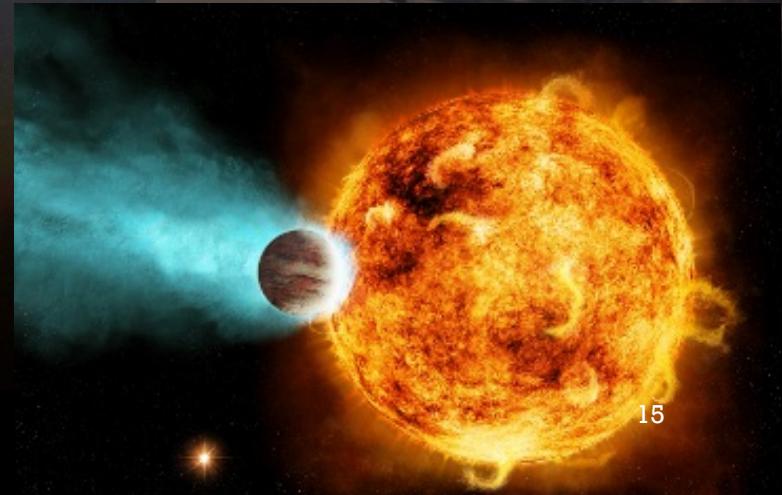
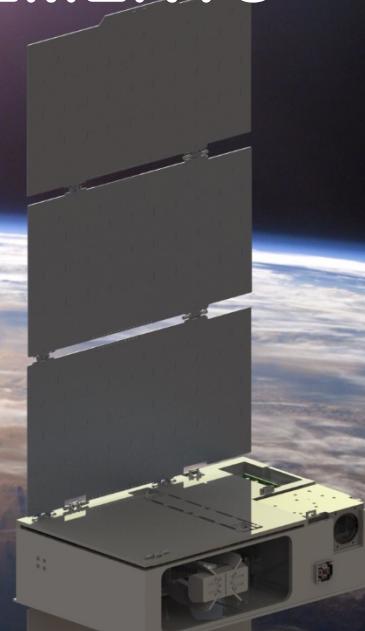
7.8 cm x 23.9 cm x 34.4 cm



CUTE: A NEW APPROACH TO ATMOSPHERIC MASS-LOSS MEASUREMENTS

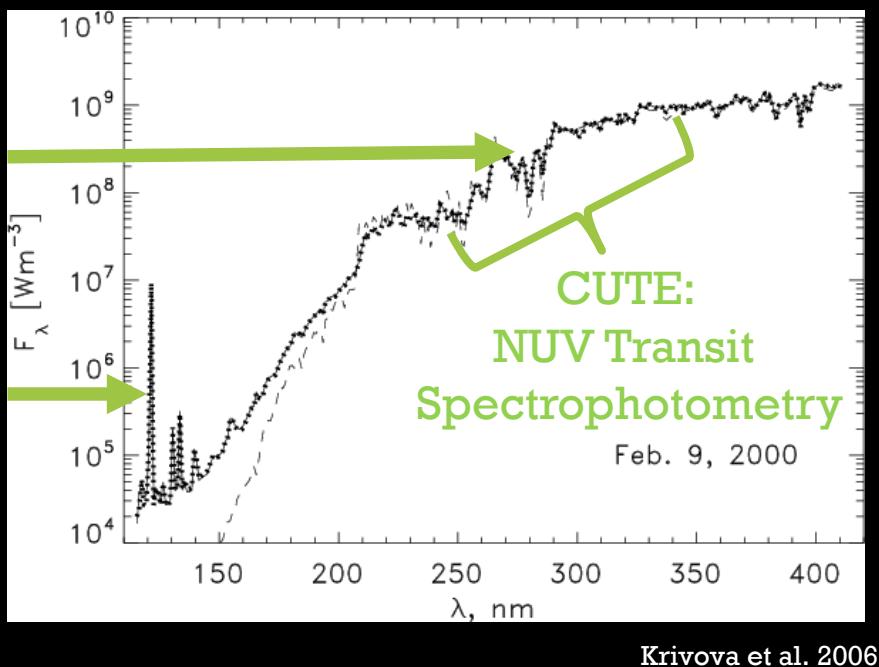
Survey of ~12-24 short-period transiting planets around nearby stars:

- 1) Atmospheric mass-loss rates
- 2) Escaping atmosphere composition



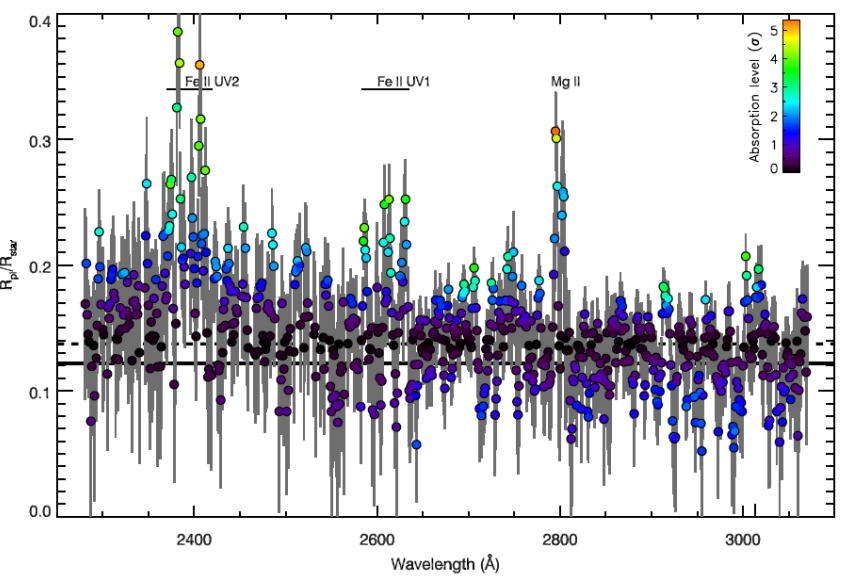
CUTE: A NEW APPROACH TO ATMOSPHERIC MASS-LOSS MEASUREMENTS

Source: SDO



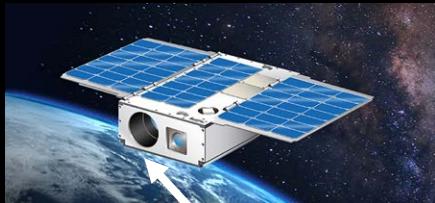
- Uncertain chromospheric intensity distribution (e.g., Llama & Shkolnik 2015).
- The NUV has both a more uniform, mainly photospheric, intensity distribution compared to FUV, and an **overall brighter background for transit observations, ~50-1000x brighter for K – F type stars.**

CUTE: A NEW APPROACH TO ATMOSPHERIC MASS-LOSS MEASUREMENTS



Sing et al. 2019

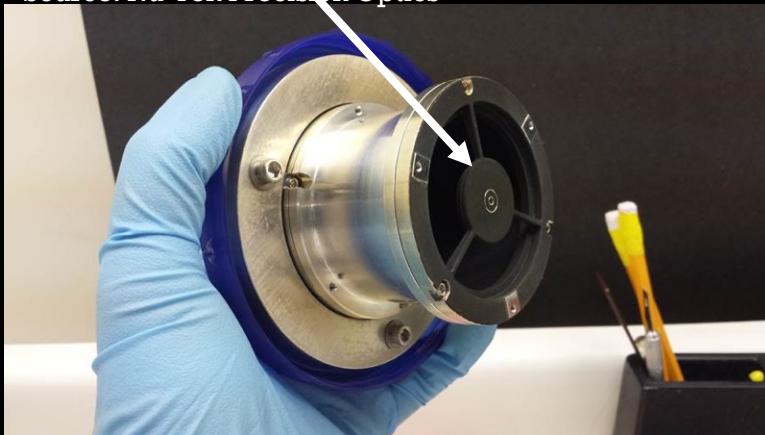
- Brighter stellar flux enables spectroscopy in a correspondingly smaller platform
- Spectroscopy required to isolate escaping gas species



CUTE Telescope

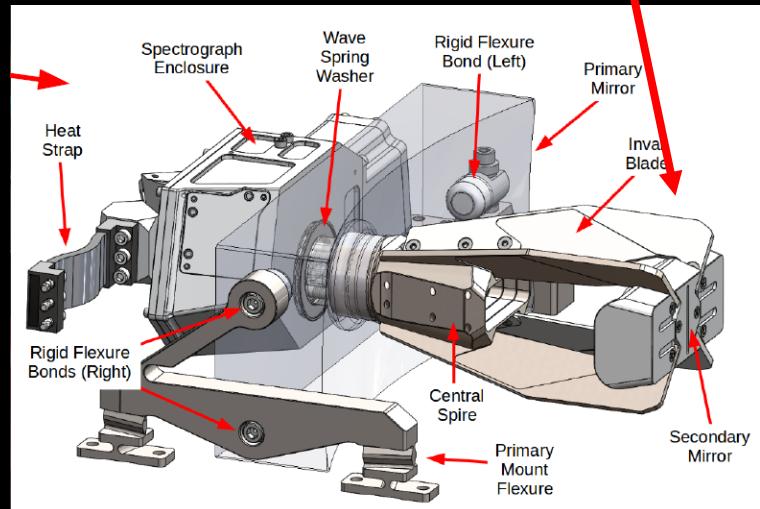


Source: Nu-Tek Precision Optics



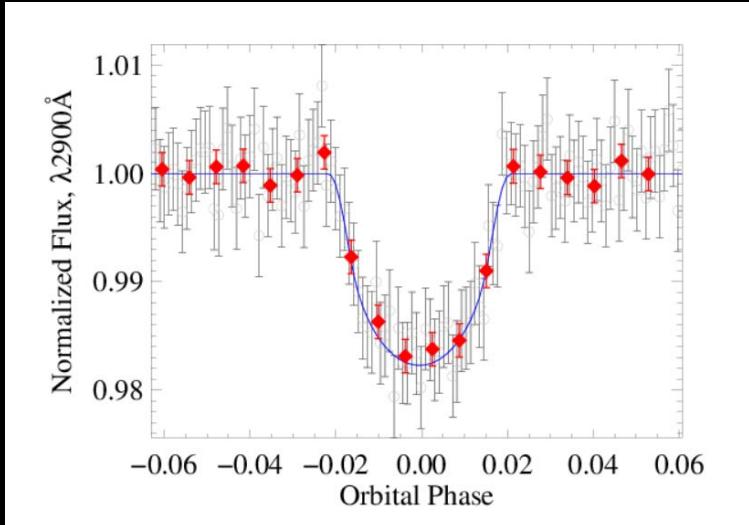
Geometric clear area for a
9cm Cassegrain: $A_T \sim 47 \text{ cm}^2$

See CUTE design overview in Fleming et al. (2018)



Geometric clear area for a 20 x 8
cm Cassegrain: $A_{\text{CUTE}} \sim 140 \text{ cm}^2$

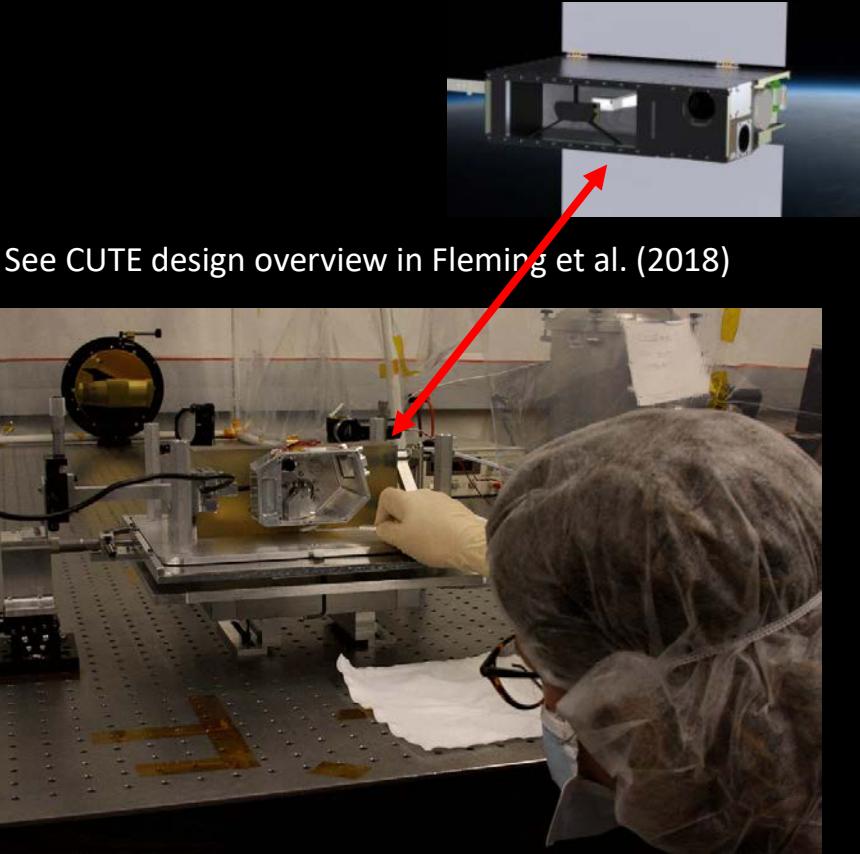
$$A_{T,r}/A_{\text{CUTE}} = 3 \times \text{more collecting area}$$



CUTE will achieve $>3\sigma$ detections of NUV continuum transits as low as 0.1% depth for the brightest targets, and < 1% for all baseline targets with 5+ lightcurves per target:

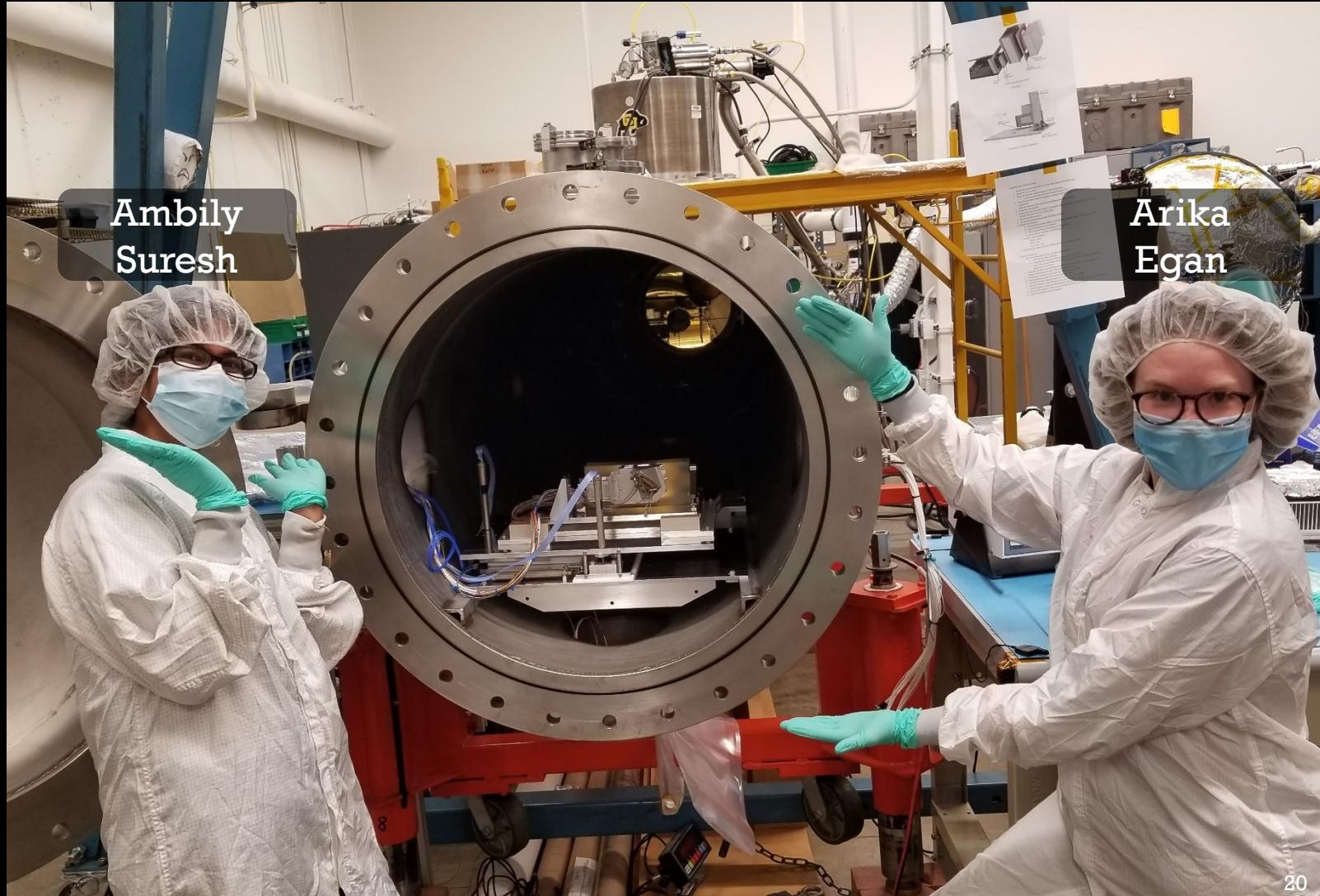
- Continuum transit sensitivity to 0.7% depth for median target over 1 transit (1.2% in Mg II)

= Capable of detecting geometric transit and atmospheric transit





CUTE Telescope Testing – Sept 2020



20

See CUTE design overview in Fleming et al. (2018); Egan et al. (2018);
mission overview in France et al. (2020) – Nature Astronomy – in press

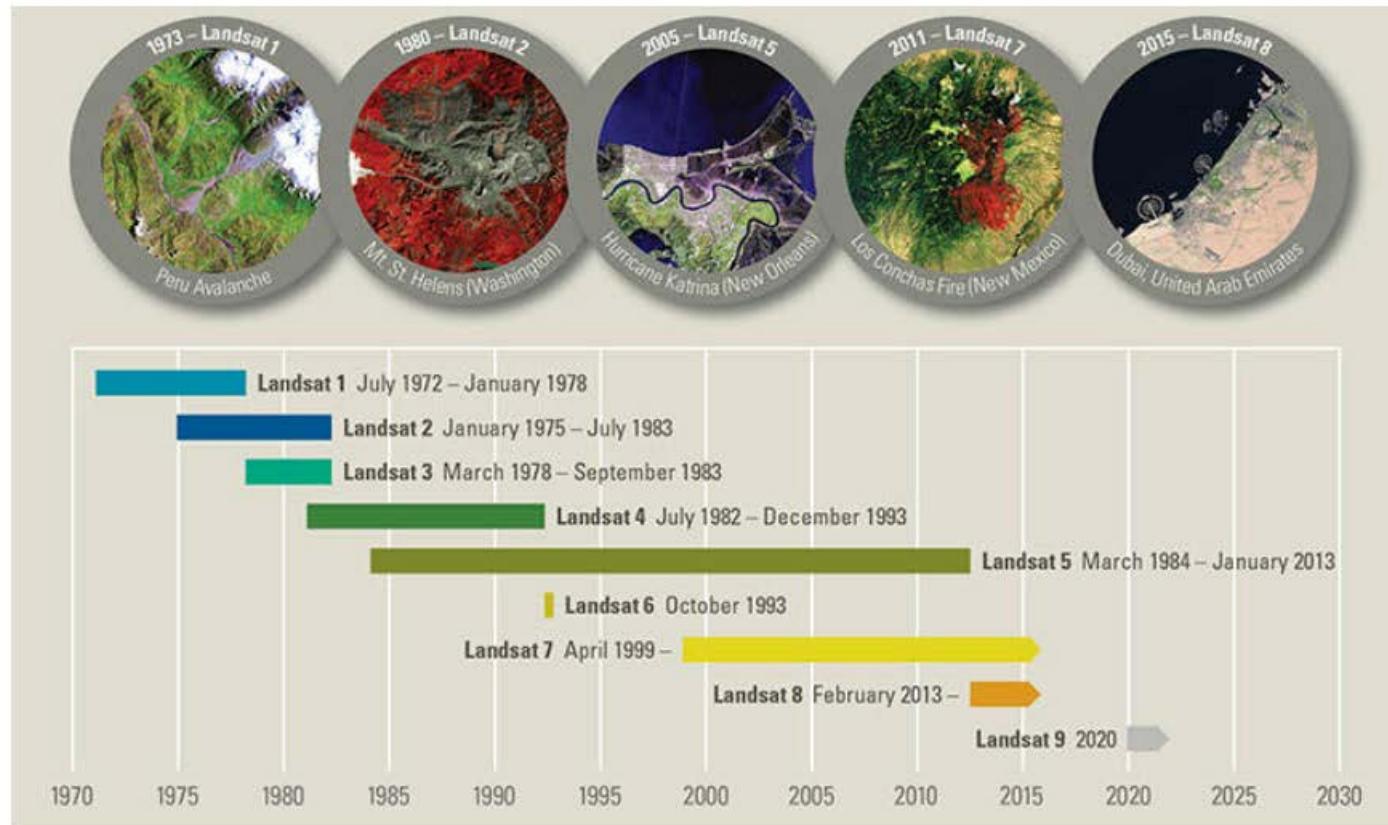
CUTE Status

When will the Landsat 9 satellite be launched?

Landsat 9—a partnership between the USGS and NASA—has a launch readiness date of December 2020.

Landsat 9 will be launched from Space Launch Complex 3E at Vandenberg Air Force Base in California and will be delivered into orbit by a United Launch Alliance Atlas V 401 launch vehicle.

Learn more: [Landsat 9 Mission](#)



CUTE

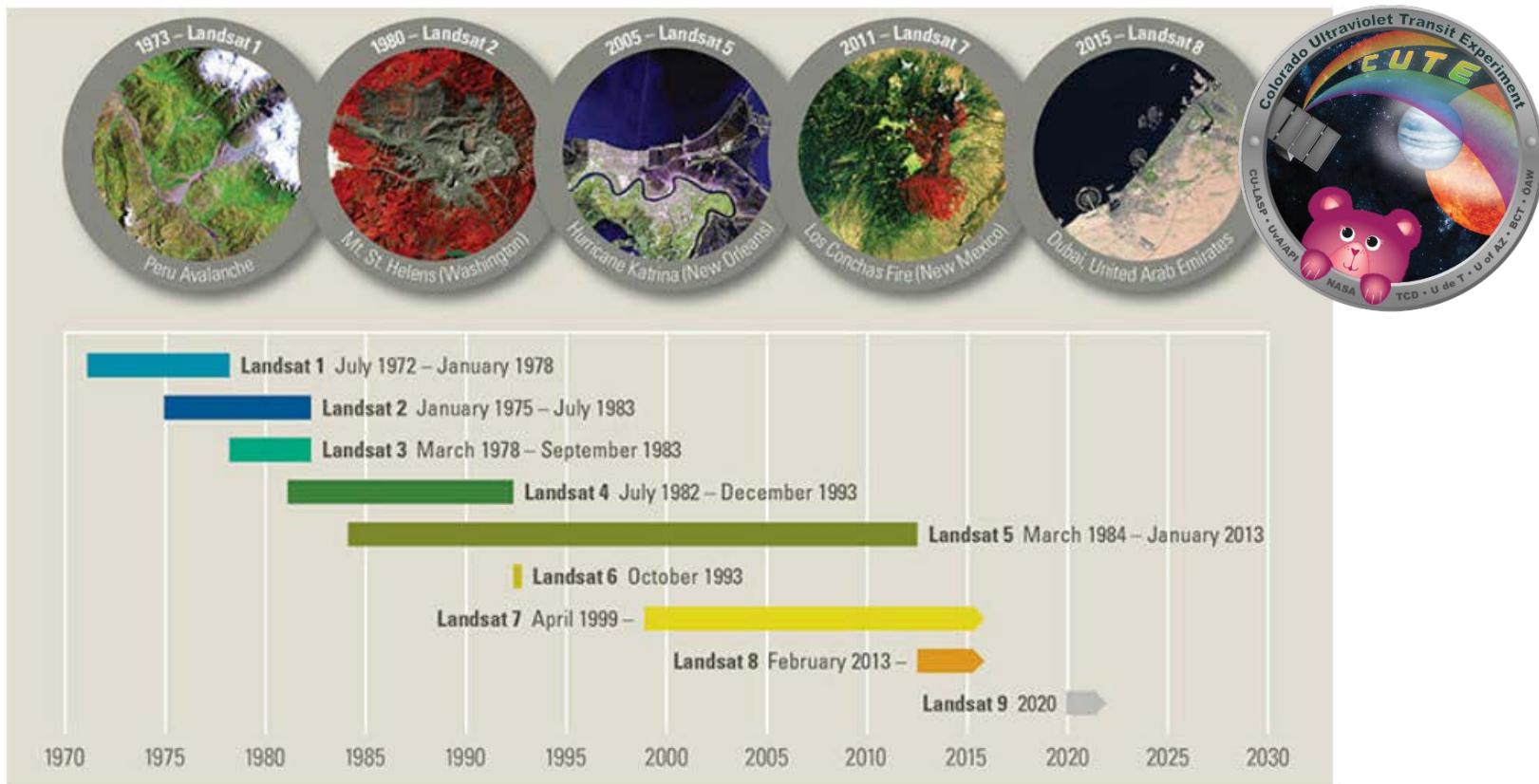
When will the Landsat 9 satellite be launched?

Landsat 9—a partnership between the USGS and NASA—has a launch readiness date of ~~December 2020~~.

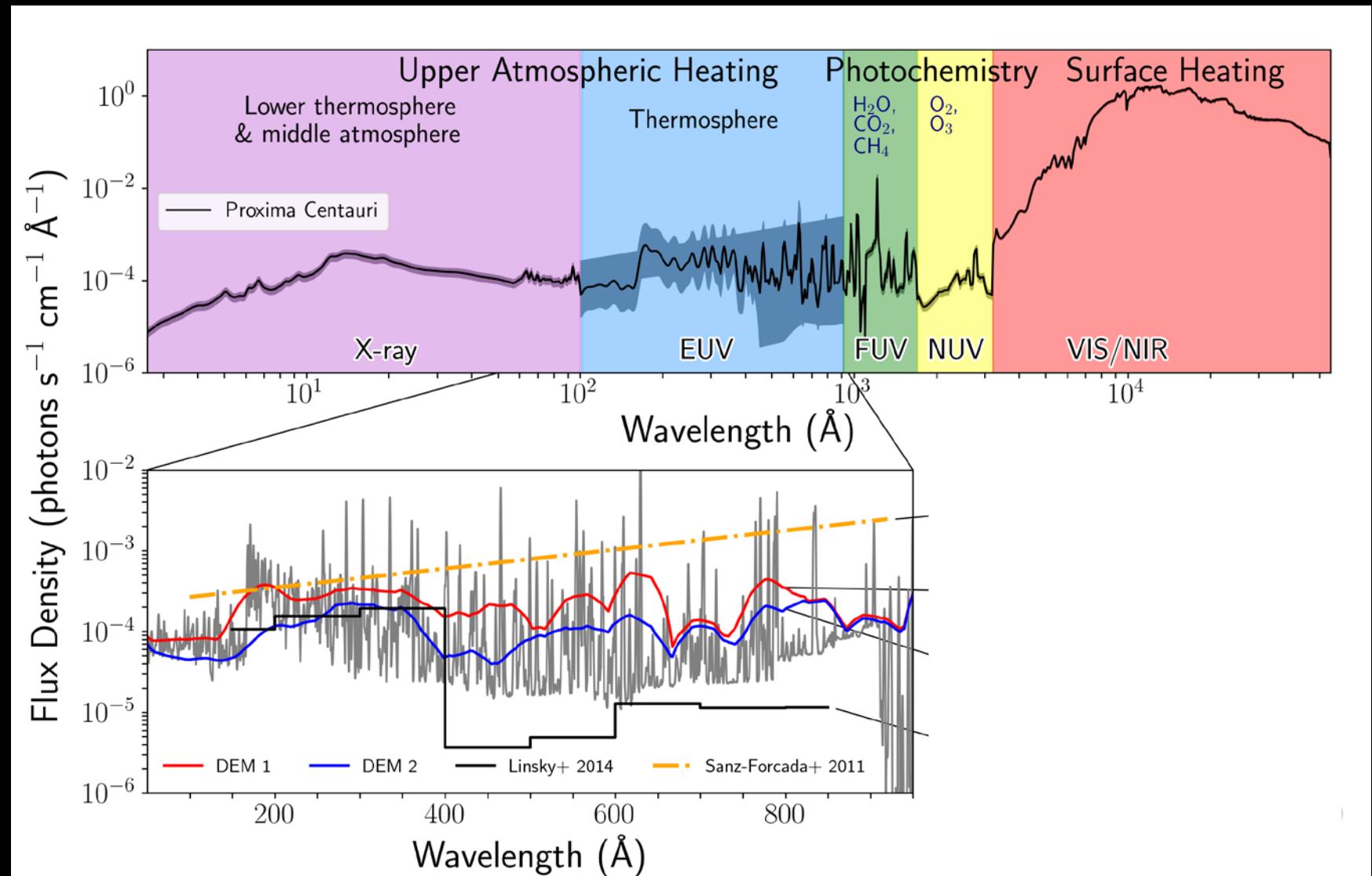
Landsat 9 will be launched from Space Launch Complex 3E at Vandenberg Air Force Base in California to be delivered into orbit by a United Launch Alliance Atlas V 401 launch vehicle.

Sept 2021

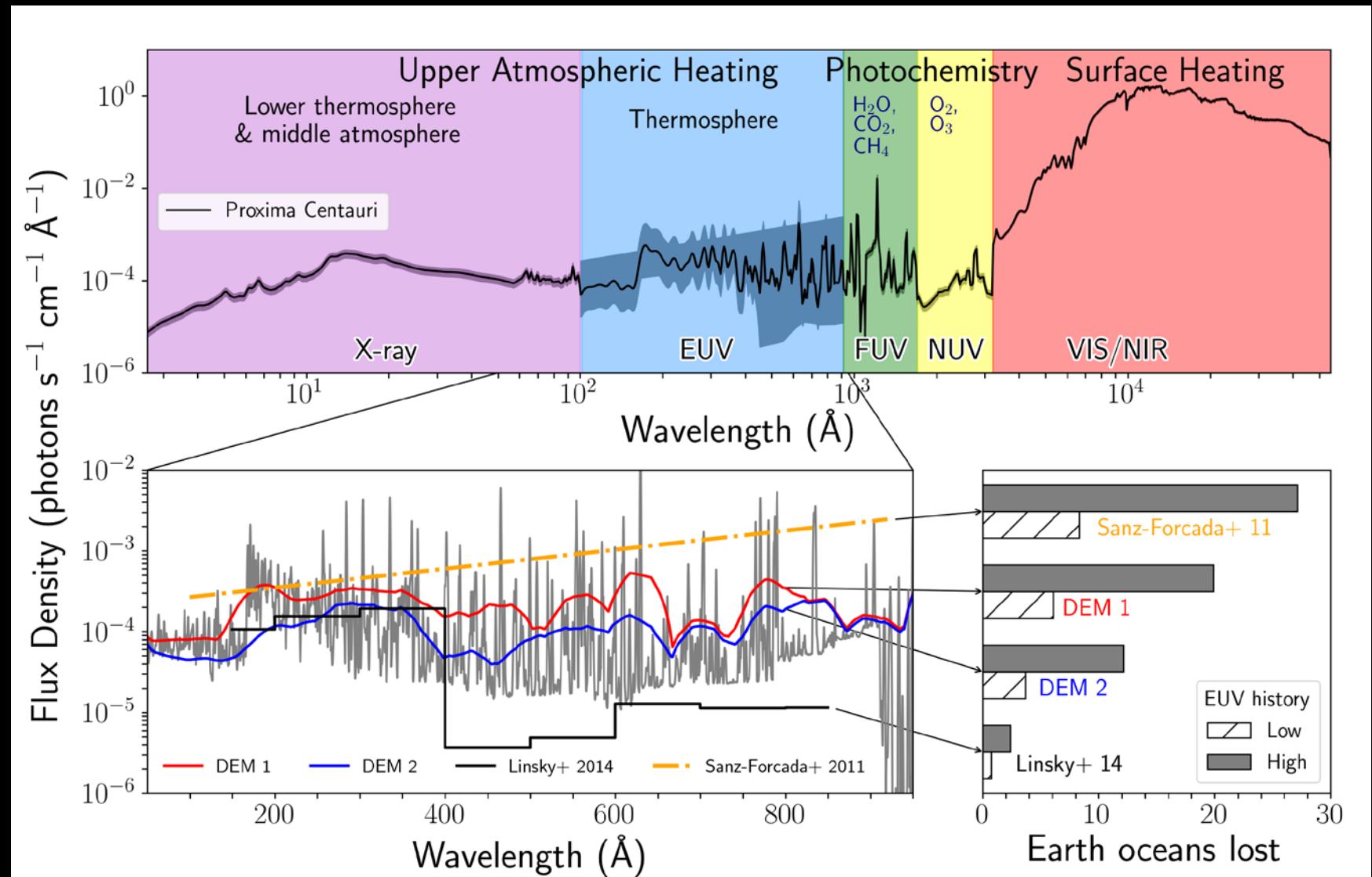
Learn more: [Landsat 9 Mission](#)



EUV environment remains a key uncertainty for all atmospheric escape calculations



EUV environment remains a key uncertainty for all atmospheric escape calculations



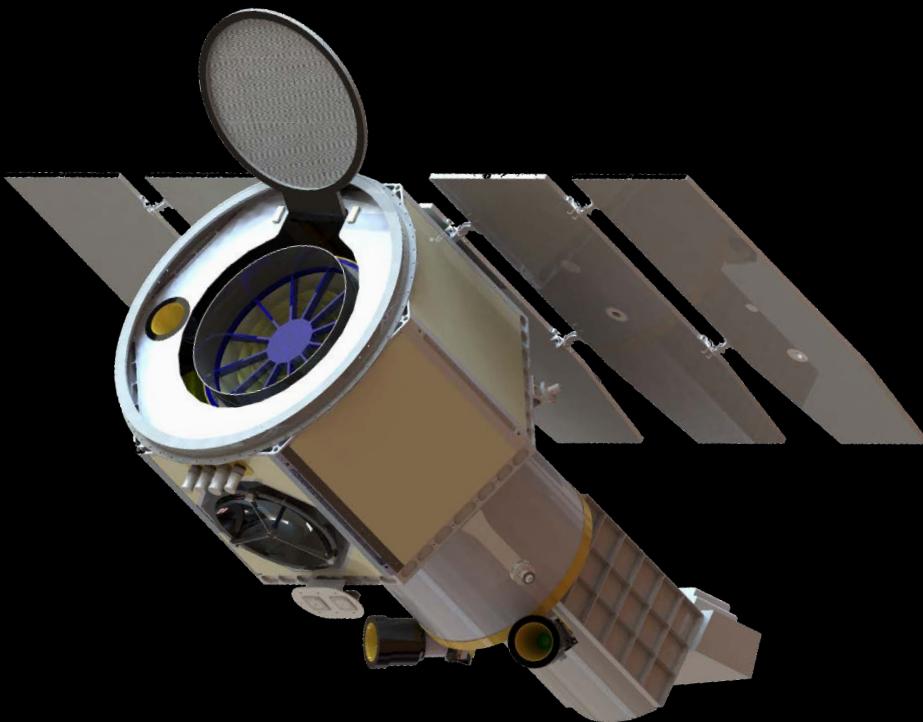


The ESCAPE Small Explorer Concept

(Euv Stellar Characterization for Atmospheric Physics and Evolution)

A NASA Small Explorer mission
currently in Phase A

PI – K. France

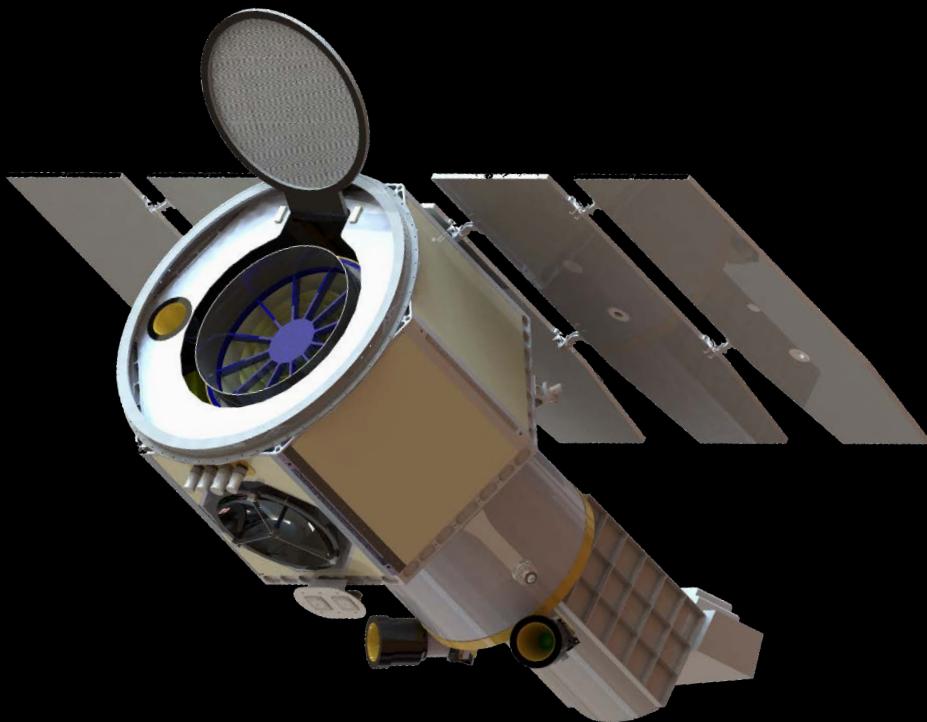




ESCAPE

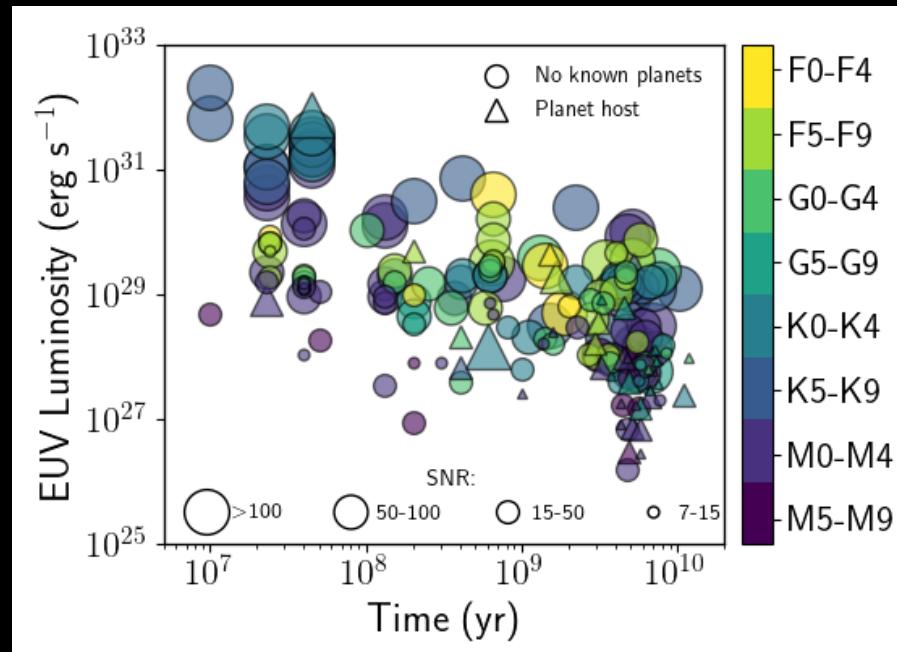
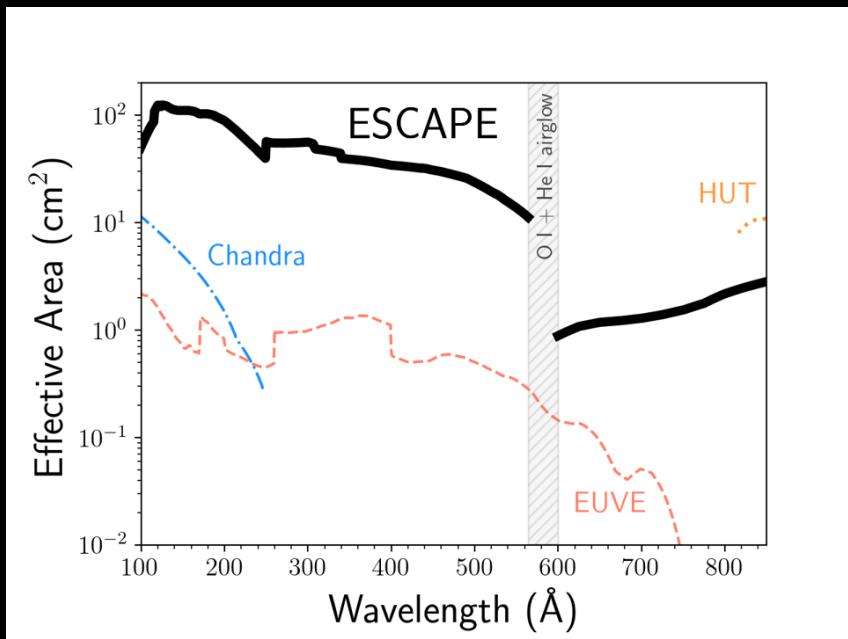
EUV & FUV (70 – 1650 Å)
spectroscopy of 200 stars,
spectral types F - M

Deep monitoring
observations of 24 targets of
interest (flare and CME
frequency)



The ESCAPE Small Explorer Concept

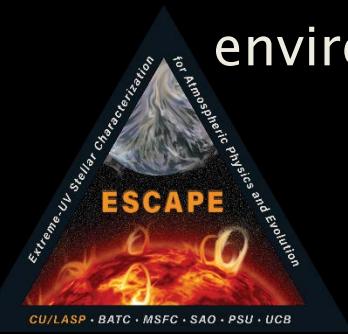
(Euv Stellar Characterization for Atmospheric Physics and Evolution)



> 100 x sensitivity of EUVE:

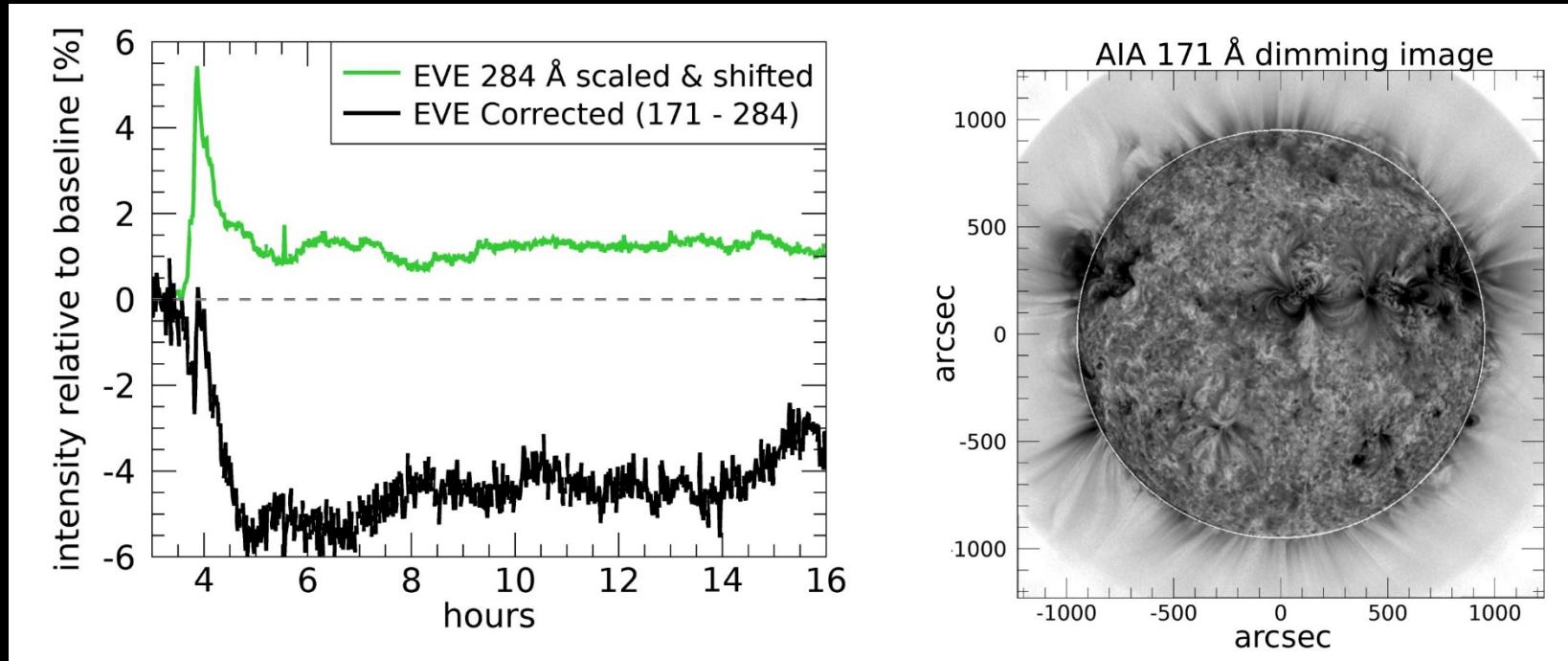
First statistical study of EUV irradiance on planetary environments, capturing important stellar/planetary timescales.

- 1) Evolutionary (Myr - Gyr)
- 2) Rotation/Stellar Cycle (days - years)
- 3) Impulsive (minutes - hours)



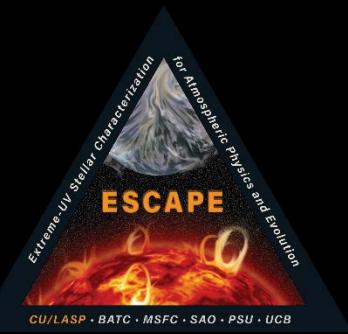
The ESCAPE Small Explorer Concept

(Euv Stellar Characterization for Atmospheric Physics and Evolution)



> 100 x sensitivity of EUVE:

- 1) CME frequency distribution via **coronal dimming** (10 – 15 F, G, and K stars)
- 2) Relationship between flares and CMEs
- 3) CME kinetic energy for brightest stars

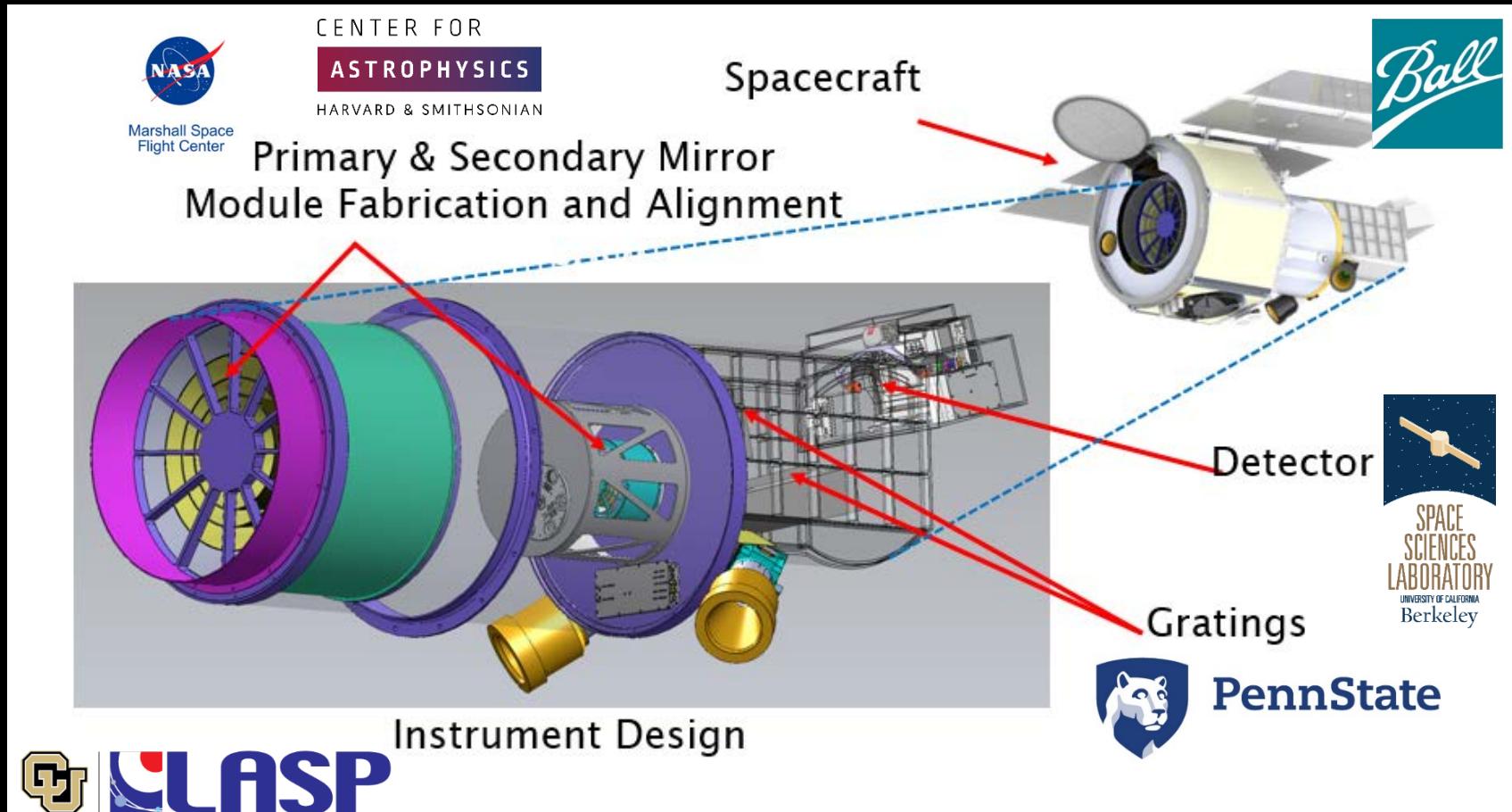


CU/LASP • BATC • MSFC • SAO • PSU • UCB

Adapted from Mason et al. (2016)

The ESCAPE Small Explorer Concept

(Euv Stellar Characterization for Atmospheric Physics and Evolution)





ESCAPE



Phase A Concept Study:
today – March 2021

Phase A Site Visit:
June 2021

Phase B down select:
October 2021 (?)

Launch:
~August 2025

Science Mission:
2025 – 2027

