

# ULTRAVIOLET ASTRONOMY IN THE XXI CENTURY



e-Workshop 2020 – October 27-29

# WSO-UV/UVSPEX for characterization of Earth-like exoplanets

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Shugarov

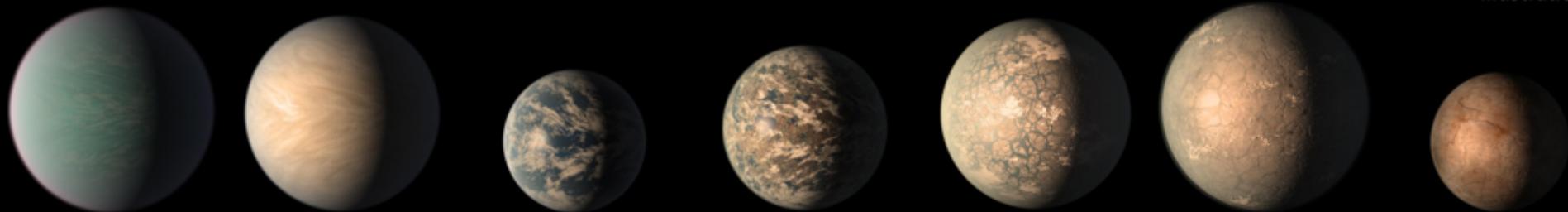
Sketches Of Spain



NASA

**TRAPPIST-1  
System**

Feb. 2018



	<b>b</b>	<b>c</b>	<b>d</b>	<b>e</b>	<b>f</b>	<b>g</b>	<b>h</b>
<i>Orbital Period</i>	<b>1.51 days</b>	<b>2.42 days</b>	<b>4.05 days</b>	<b>6.10 days</b>	<b>9.21 days</b>	<b>12.36 days</b>	<b>18.76 days</b>
<i>Distance to Star</i>	<b>0.0115 AU</b>	<b>0.0158 AU</b>	<b>0.0223 AU</b>	<b>0.0293 AU</b>	<b>0.0385 AU</b>	<b>0.0469 AU</b>	<b>0.0619 AU</b>
<i>Planet Radius</i>	<b><math>1.12 R_{\text{earth}}</math></b>	<b><math>1.10 R_{\text{earth}}</math></b>	<b><math>0.78 R_{\text{earth}}</math></b>	<b><math>0.91 R_{\text{earth}}</math></b>	<b><math>1.05 R_{\text{earth}}</math></b>	<b><math>1.15 R_{\text{earth}}</math></b>	<b><math>0.77 R_{\text{earth}}</math></b>
<i>Planet Mass</i>	<b><math>1.02 M_{\text{earth}}</math></b>	<b><math>1.16 M_{\text{earth}}</math></b>	<b><math>0.30 M_{\text{earth}}</math></b>	<b><math>0.77 M_{\text{earth}}</math></b>	<b><math>0.93 M_{\text{earth}}</math></b>	<b><math>1.15 M_{\text{earth}}</math></b>	<b><math>0.33 M_{\text{earth}}</math></b>
<i>Planet Density</i>	<b><math>0.73 \rho_{\text{earth}}</math></b>	<b><math>0.88 \rho_{\text{earth}}</math></b>	<b><math>0.62 \rho_{\text{earth}}</math></b>	<b><math>1.02 \rho_{\text{earth}}</math></b>	<b><math>0.82 \rho_{\text{earth}}</math></b>	<b><math>0.76 \rho_{\text{earth}}</math></b>	<b><math>0.72 \rho_{\text{earth}}</math></b>
<i>Surface Gravity</i>	<b><math>0.81 g</math></b>	<b><math>0.96 g</math></b>	<b><math>0.48 g</math></b>	<b><math>0.93 g</math></b>	<b><math>0.85 g</math></b>	<b><math>0.87 g</math></b>	<b><math>0.55 g</math></b>

**Solar System  
Rocky Planets**

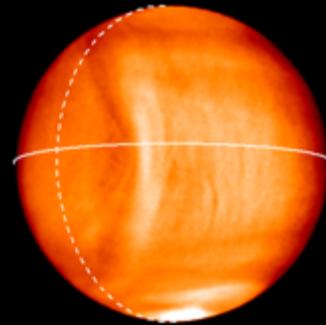

	<b>Mercury</b>	<b>Venus</b>	<b>Earth</b>	<b>Mars</b>
<i>Orbital Period</i>	<b>87.97 days</b>	<b>224.70 days</b>	<b>365.26 days</b>	<b>686.98 days</b>
<i>Distance to Star</i>	<b>0.387 AU</b>	<b>0.723 AU</b>	<b>1.000 AU</b>	<b>1.524 AU</b>
<i>Planet Radius</i>	<b><math>0.38 R_{\text{earth}}</math></b>	<b><math>0.95 R_{\text{earth}}</math></b>	<b><math>1.00 R_{\text{earth}}</math></b>	<b><math>0.53 R_{\text{earth}}</math></b>
<i>Planet Mass</i>	<b><math>0.06 M_{\text{earth}}</math></b>	<b><math>0.82 M_{\text{earth}}</math></b>	<b><math>1.00 M_{\text{earth}}</math></b>	<b><math>0.11 M_{\text{earth}}</math></b>
<i>Planet Density</i>	<b><math>0.98 \rho_{\text{earth}}</math></b>	<b><math>0.95 \rho_{\text{earth}}</math></b>	<b><math>1.00 \rho_{\text{earth}}</math></b>	<b><math>0.71 \rho_{\text{earth}}</math></b>
<i>Surface Gravity</i>	<b><math>0.38 g</math></b>	<b><math>0.90 g</math></b>	<b><math>1.00 g</math></b>	<b><math>0.38 g</math></b>

Detection  
- Transit, RV



Characterization  
=>Atmosphere

# Earth and Venus



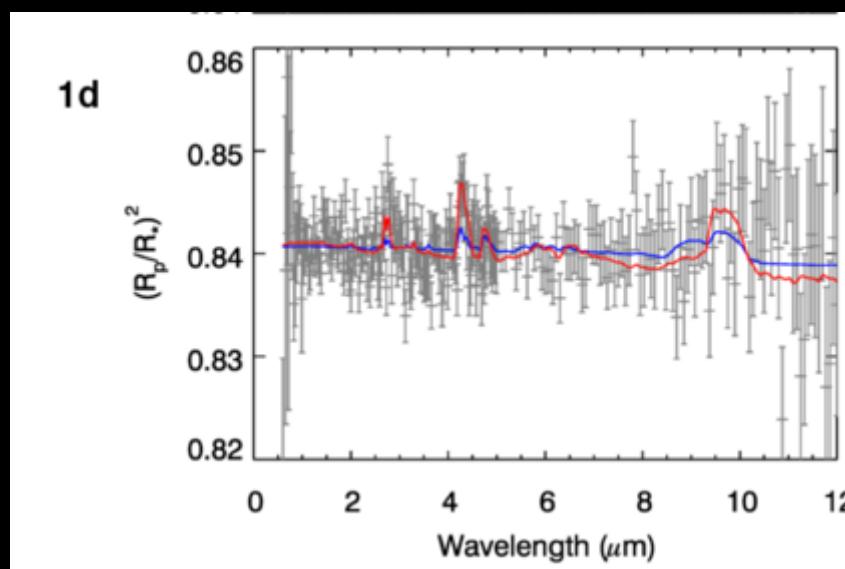
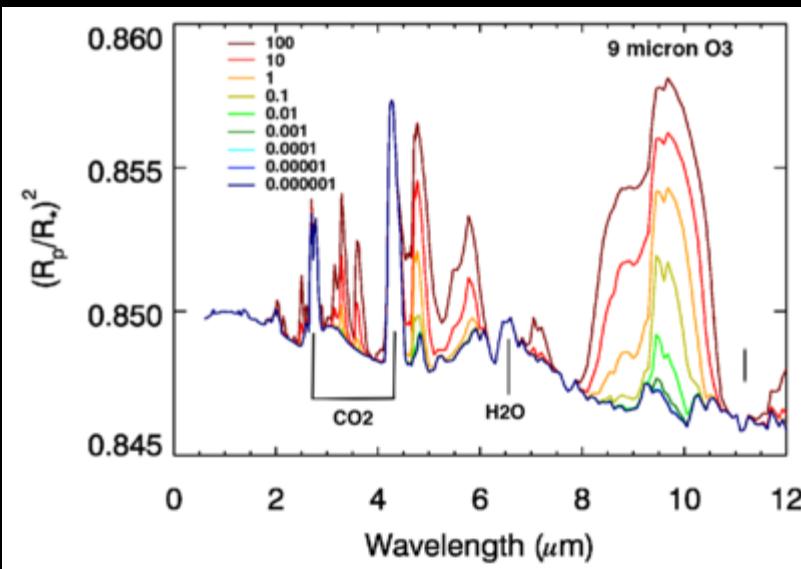
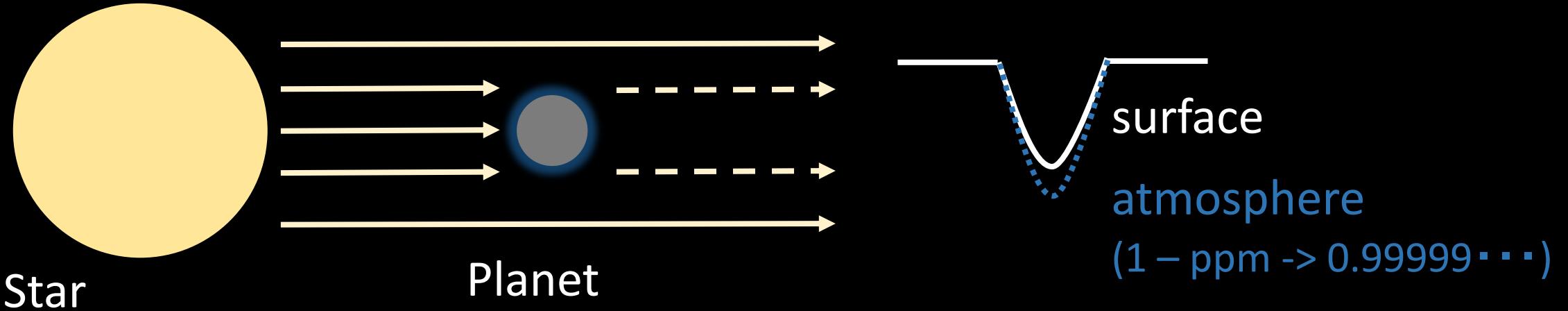
Almost the same in size  
Earth R: 6371 km  
Venus R: 6052 km

Life (Civilization)  
Ocean  
 $N_2$ (78%),  $O_2$ (21%), Low  $CO_2$   
Magnetic field

No life  
No ocean  
 $CO_2$  (96%)  
No magnetic field

How to distinguish?

# Transit spectroscopy of exoplanetary atmosphere



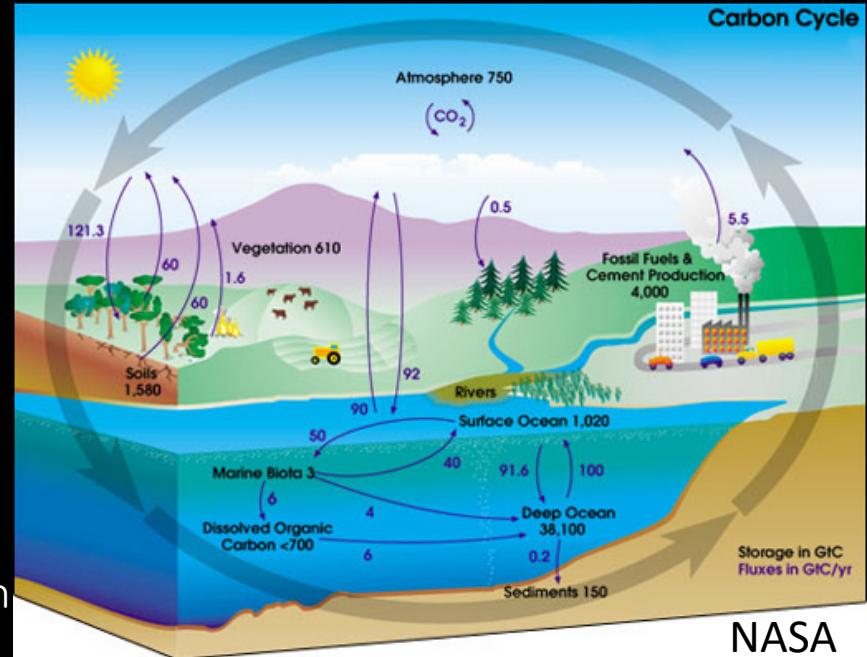
Ozone in Earth-like  
TRAPPIST-1d  
=>  
60 transits by JWST  
(~6 years)

Barstow & Irwin, 2016

# Earth and Venus



Almost the same in  
Earth R: 6371 km  
Venus R: 6052 km



Life (Civilization)  
Ocean  
N<sub>2</sub>(78%), O<sub>2</sub>(21%), Low CO<sub>2</sub>  
Magnetic field

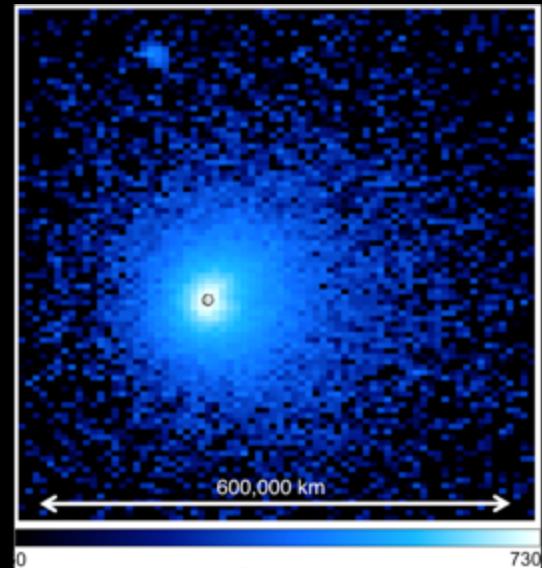
No life  
No ocean  
CO<sub>2</sub> (96%)  
No magnetic field

How to distinguish?

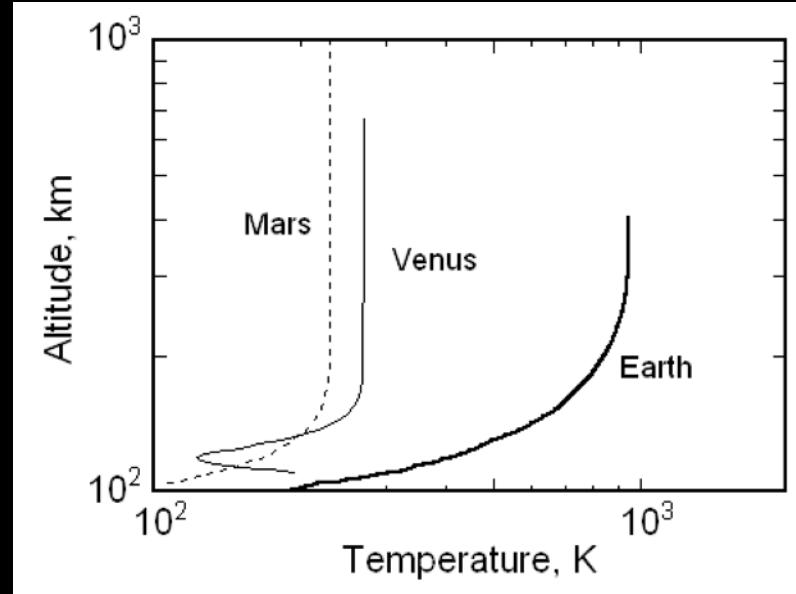
# Upper atmosphere (Exosphere)

$\text{CO}_2$  is a coolant for upper atmosphere.  
Earth (Low  $\text{CO}_2$ )  $\rightarrow$  1000 K  
Venus, Mars ( $\text{CO}_2$  dominant)  $\rightarrow$  200-300 K

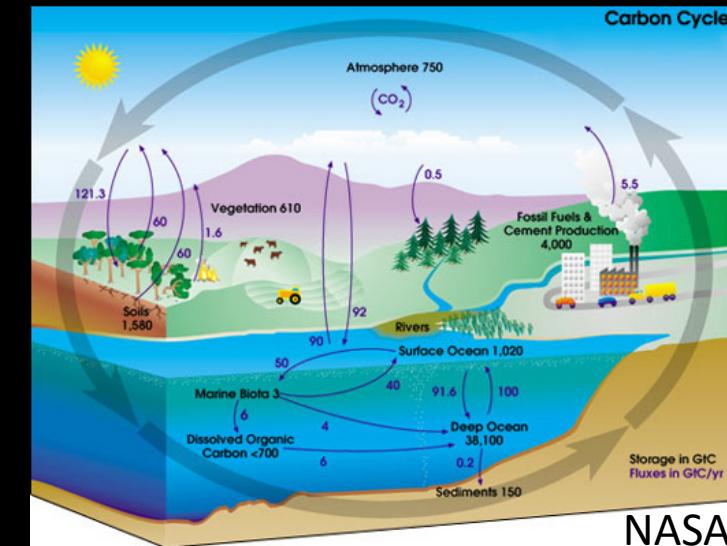
$\text{CO}_2$  on exoplanets: detectable  
 $\text{CO}_2$  mixing ratio: difficult (Venus  $\leftarrow$  In-situ)  
(temperature profile, height of cloud, etc.)



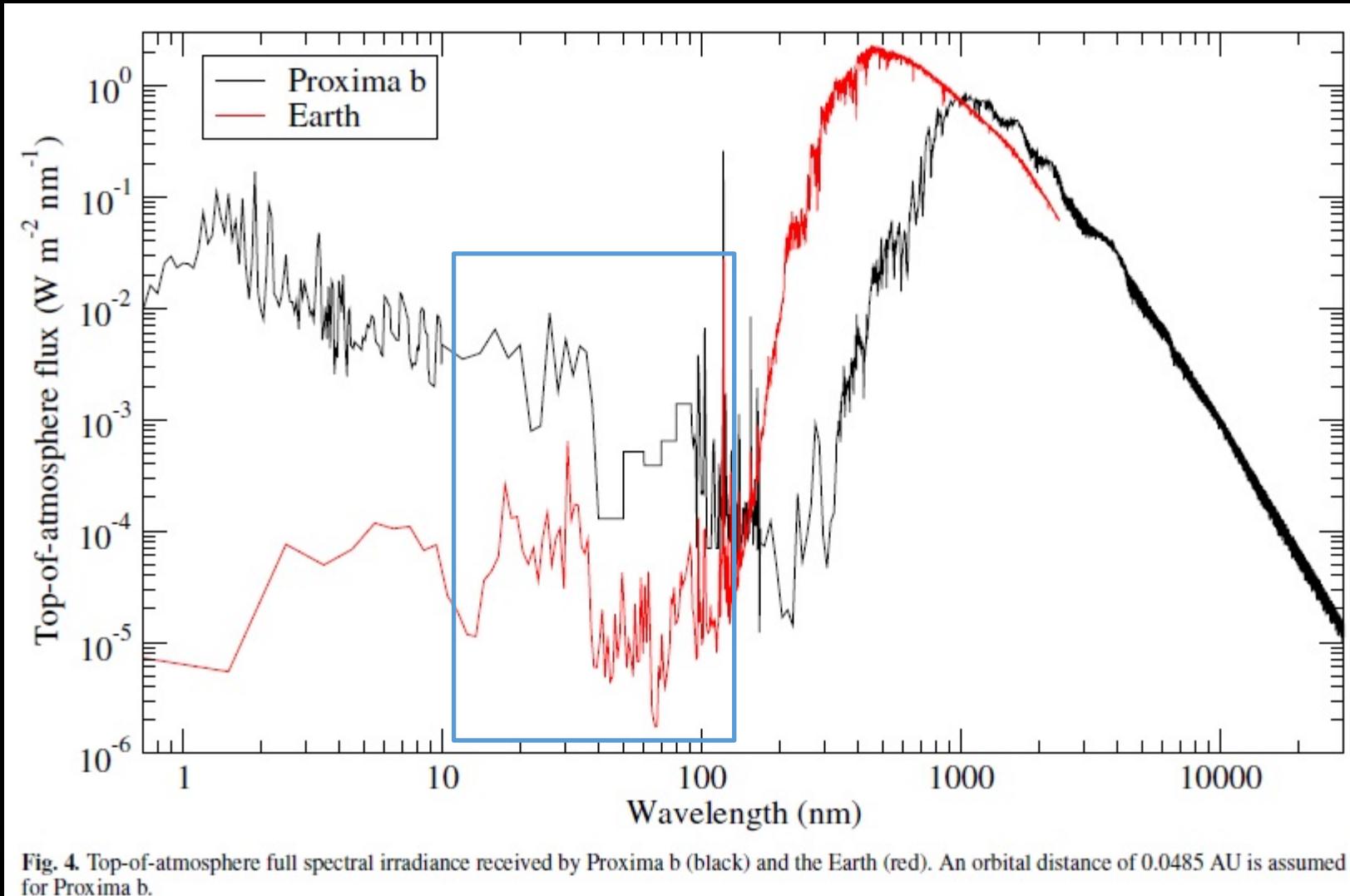
Earth's extended exosphere  
Kameda+17



Temperature of upper atmosphere  
 $\leftrightarrow$ ?  $\text{CO}_2$  mixing ratio  
 $\leftrightarrow$ ? Ocean and tectonics



# Stellar Flux at Hz of low-temperature star



Red: Solar flux (1AU)  
Black : Proxima  
Centauri' s flux at  
Proxima b (0.0485AU)

Heat source of  
Upper atmosphere  
EUV:10-105nm

- Difference between Solar and M-type stellar fluxes. (Ribas et al., 2017)

# Extended oxygen exosphere at high EUV

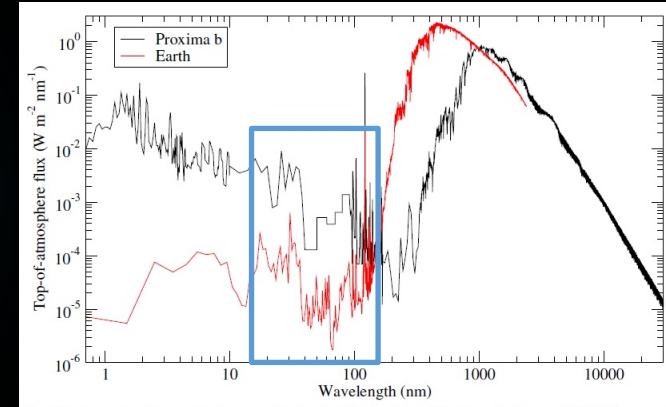
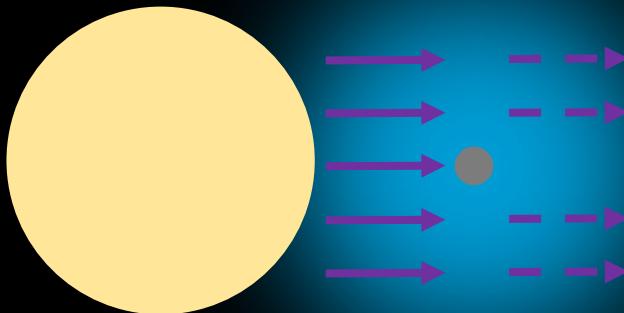
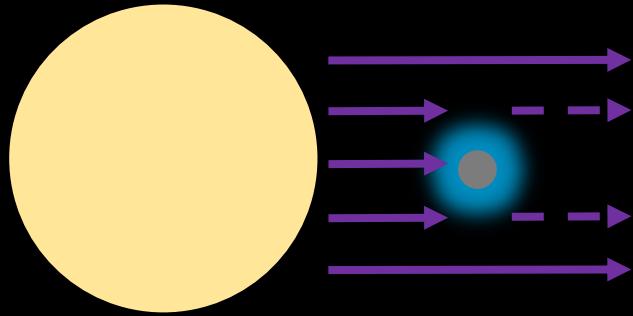
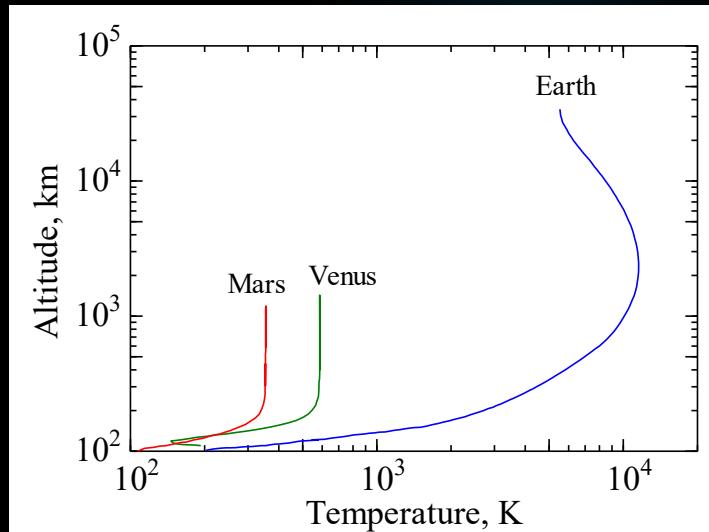
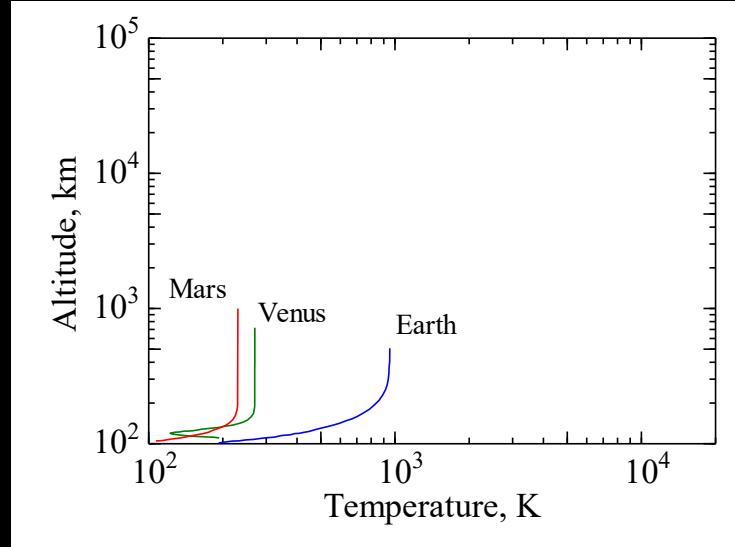


Fig. 4. Top-of-atmosphere full spectral irradiance received by Proxima b (black) and the Earth (red). An orbital distance of 0.0485 AU is assumed for Proxima b.

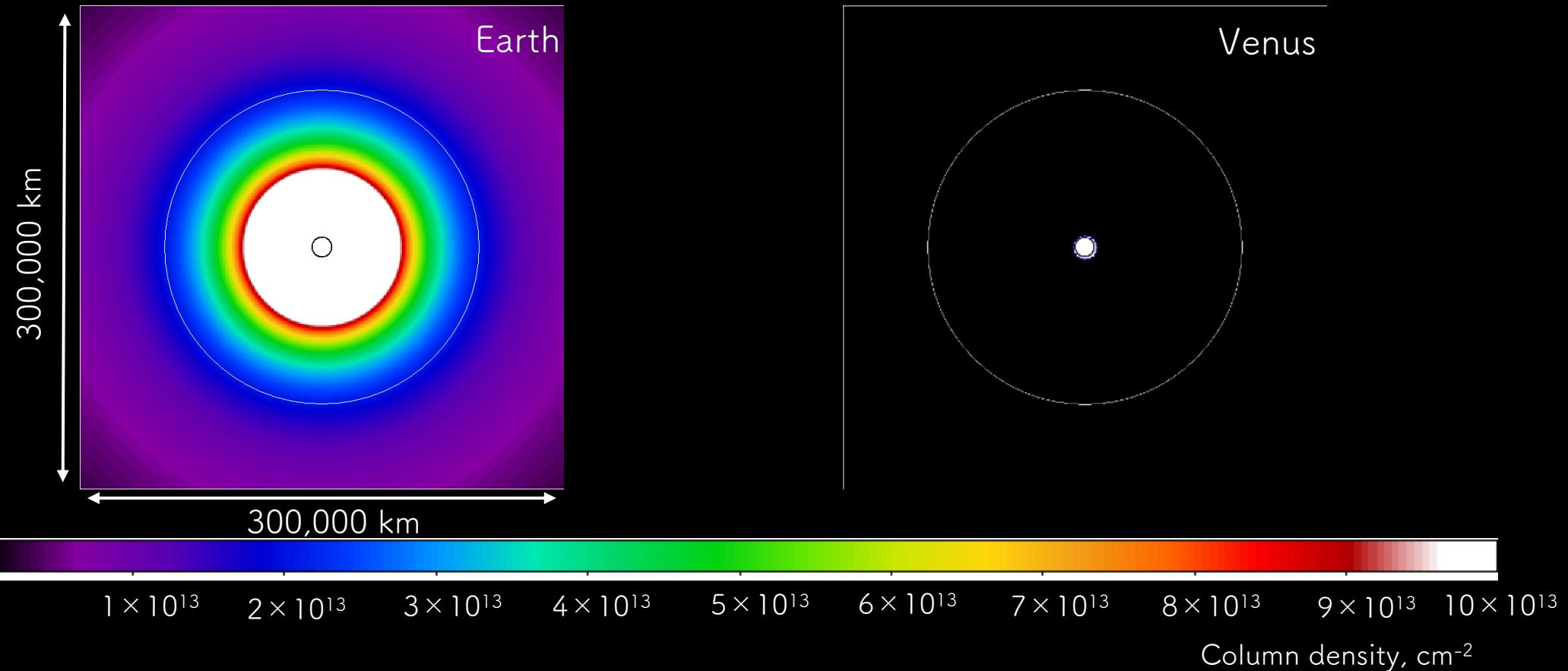


High EUV  $\rightarrow$  expanded upper atmosphere

Kulikov+07

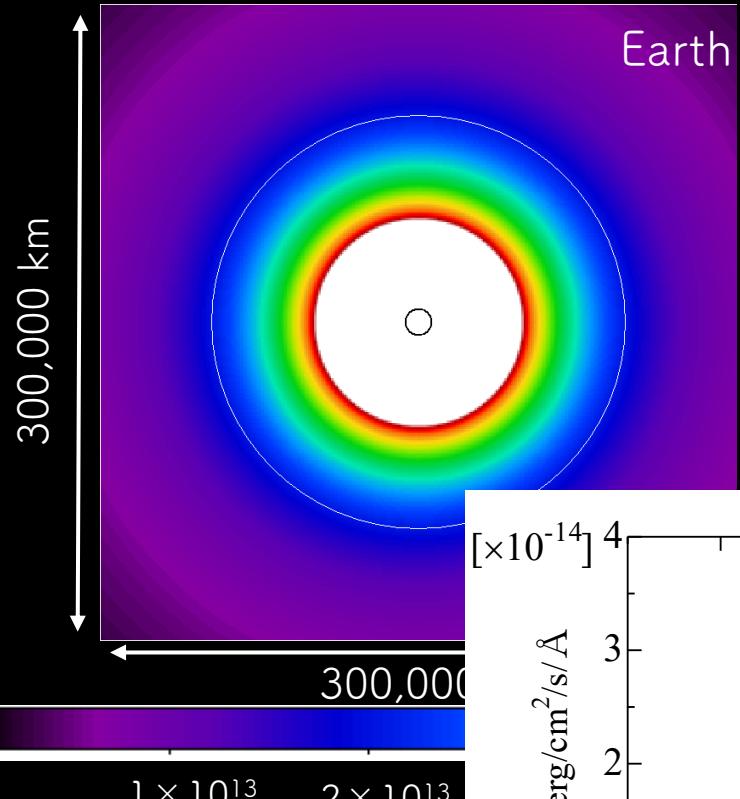
Tian+08

# Oxygen Exosphere(10EUV)

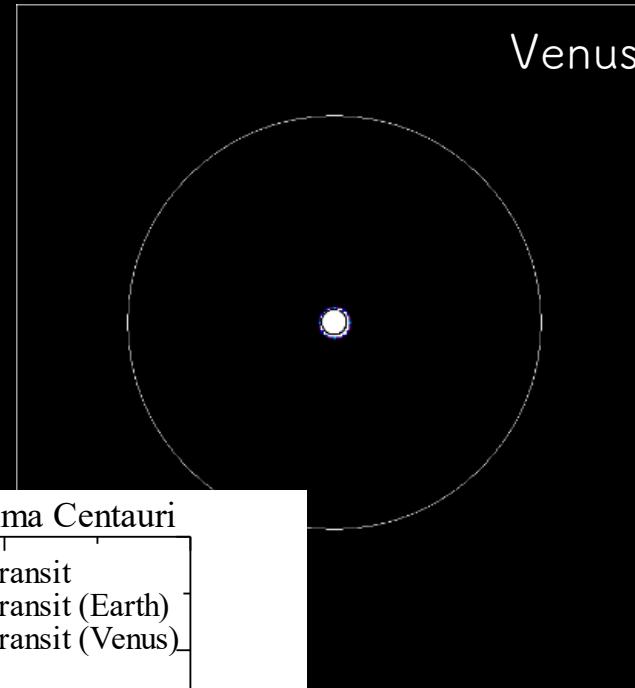
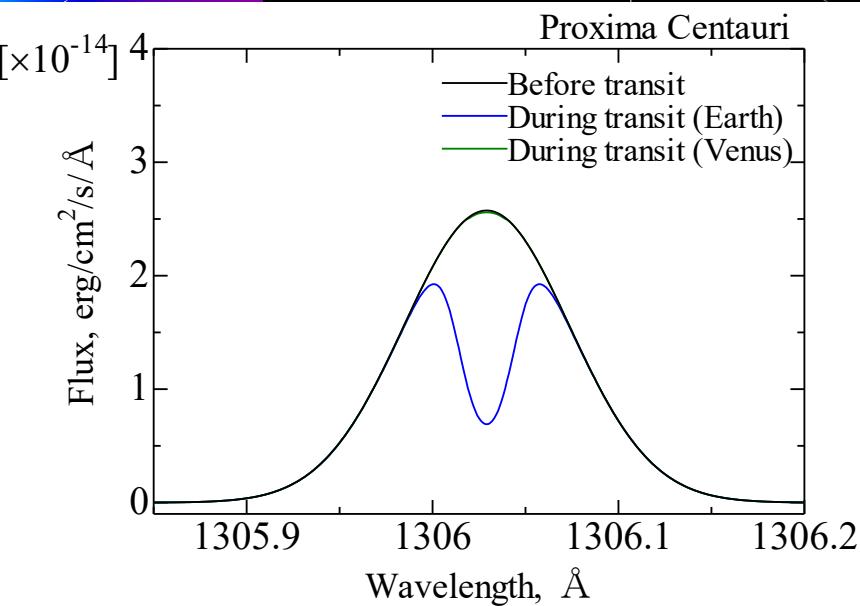


- Black circle: planet    White circle: star(Proxima Centauri)
- Earth-like case: Optically thick oxygen to  $\sim 8R_e$  ( $1 \times 10^{14} \text{ cm}^{-2} \sim \tau = 1$ )

# Oxygen Exosphere(10EUV)



- Black circle: plane
- Earth-like case: O



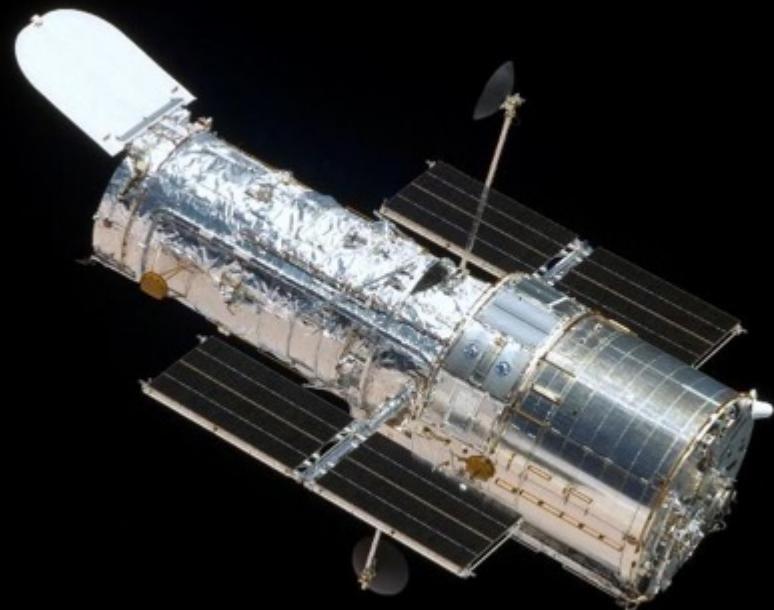
Transit depth: ~24%  
(~70% at line center)

Tavrov, Kameda+2018  
arXiv:1809.03441

$\times 10^{14} \text{ cm}^{-2} \sim \tau = 1$

Large UV space telescope!

(HST is inside Earth's oxygen corona)



[NASA]



# World Space Observatory UV (WSO-UV)



Exoplanet  
UV irradiation  
Upper atmosphere

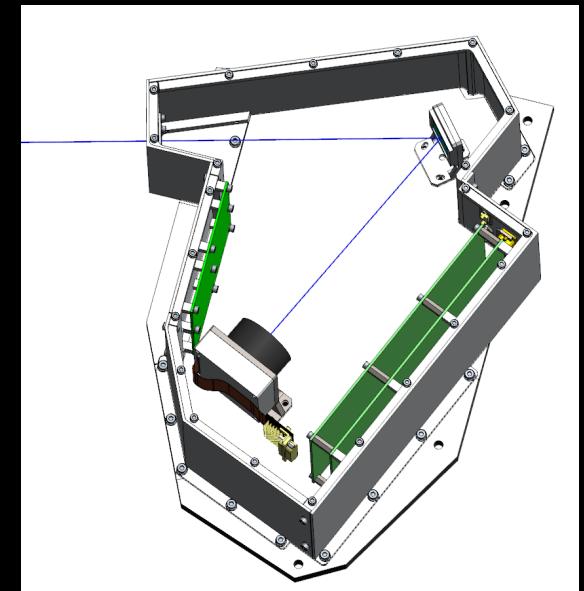
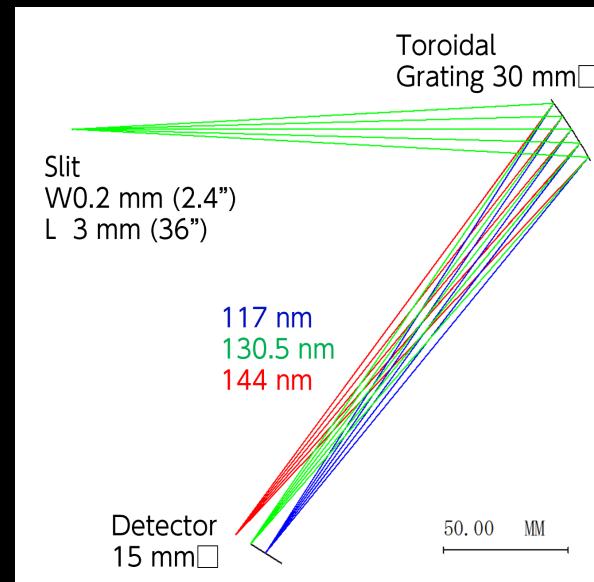
Russia, Spain (UV imager)  
+Japan (UV spectrometer)

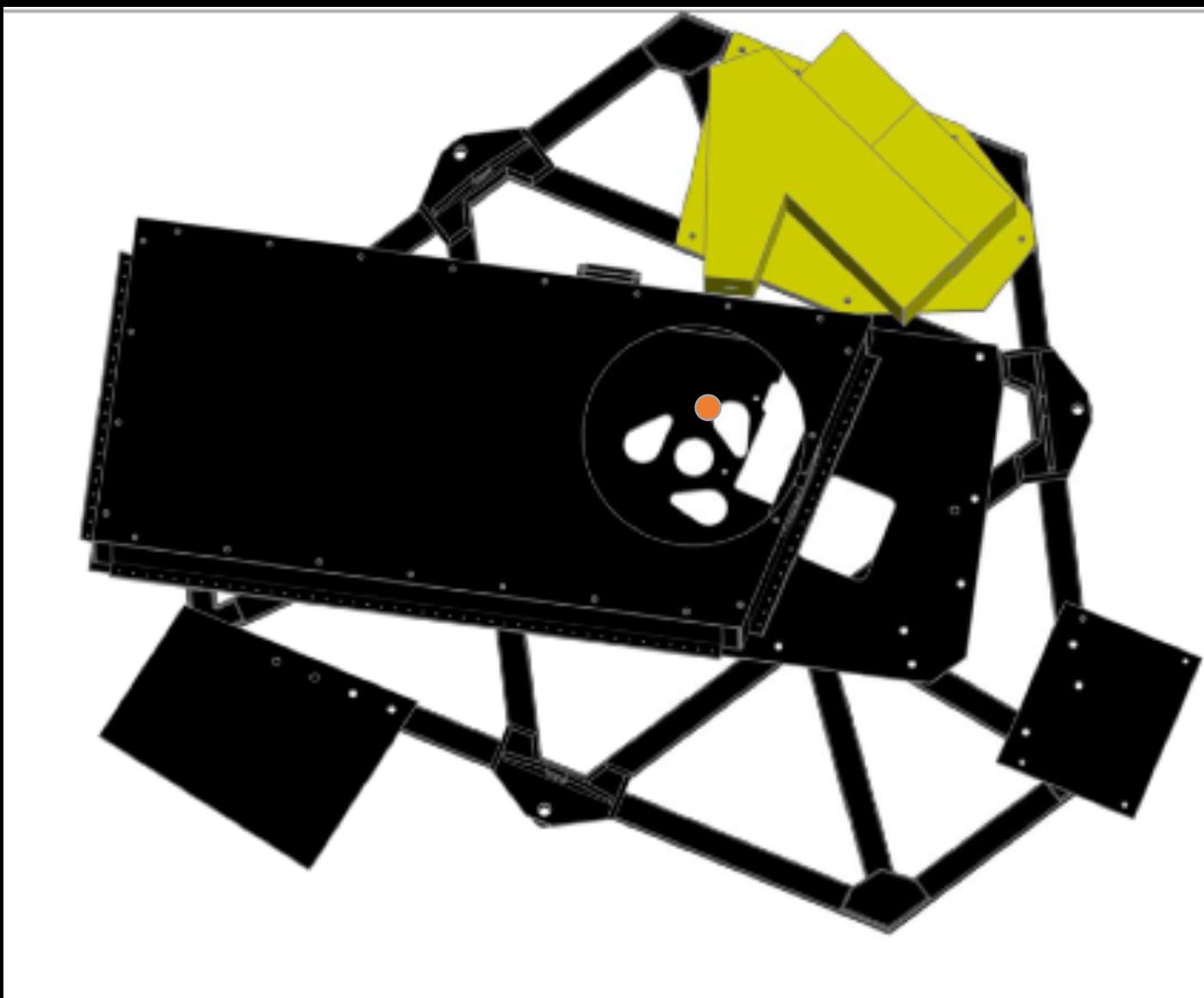
- Target launch: 2025 Oct
- Diameter : 1.7 m
- Spectral range : ~110-320 nm
- Orbit: GSO (6.6 Re) -> *outside of Oxygen geocorona* (HST ~400km)

# UV spectrograph for Exoplanets (UVSPEX)

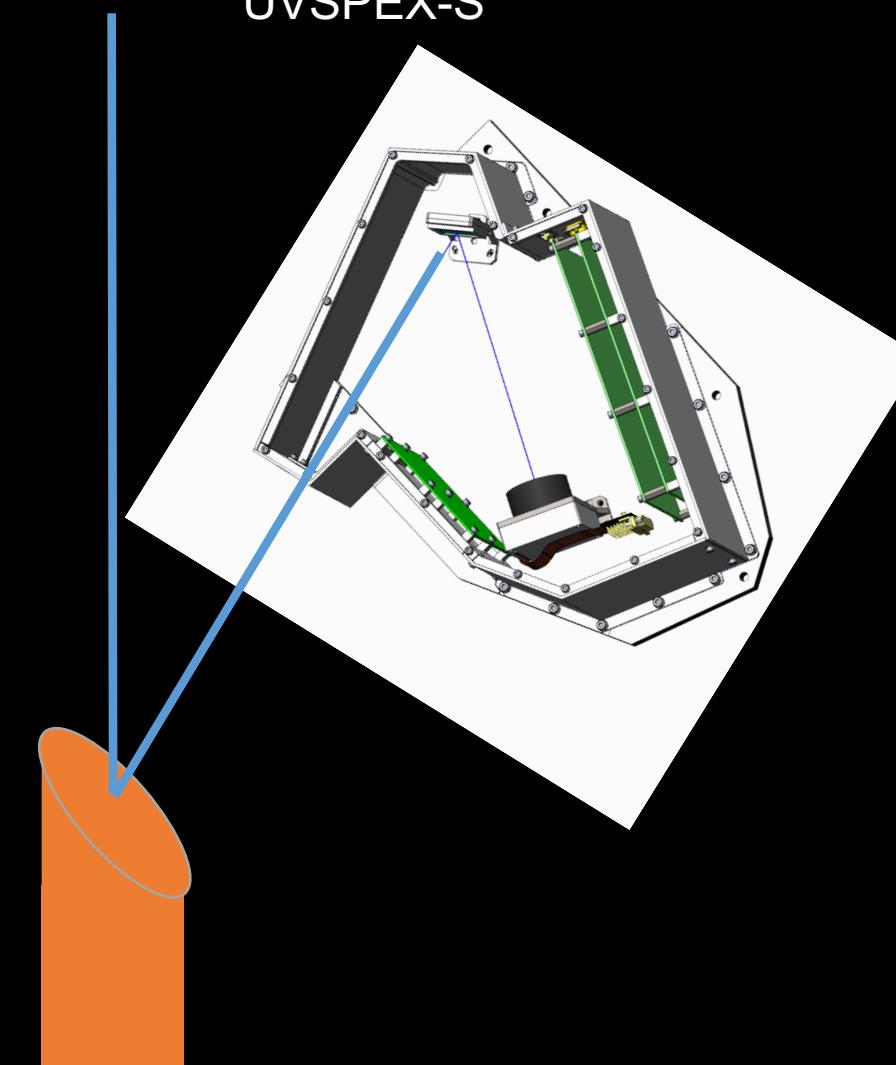
- Slit + Concave (Toroidal) grating + Detector (MCP)
- Spectral range : 117-144 nm
- Spectral resolution : ~0.3 nm @130 nm
- Slit : 2.5" (= 200 um)
- Grating : Toroidal blazed grating, Al+MgF<sub>2</sub> coating,  $\phi$  30 mm, 2400 gr/mm, f = 250 mm
- Detector : Image Intensifier (CsI photocathode + Funnel-type Microchannel Plate (MCP))
- Mass : < 7 kg
- Power: < 7 W

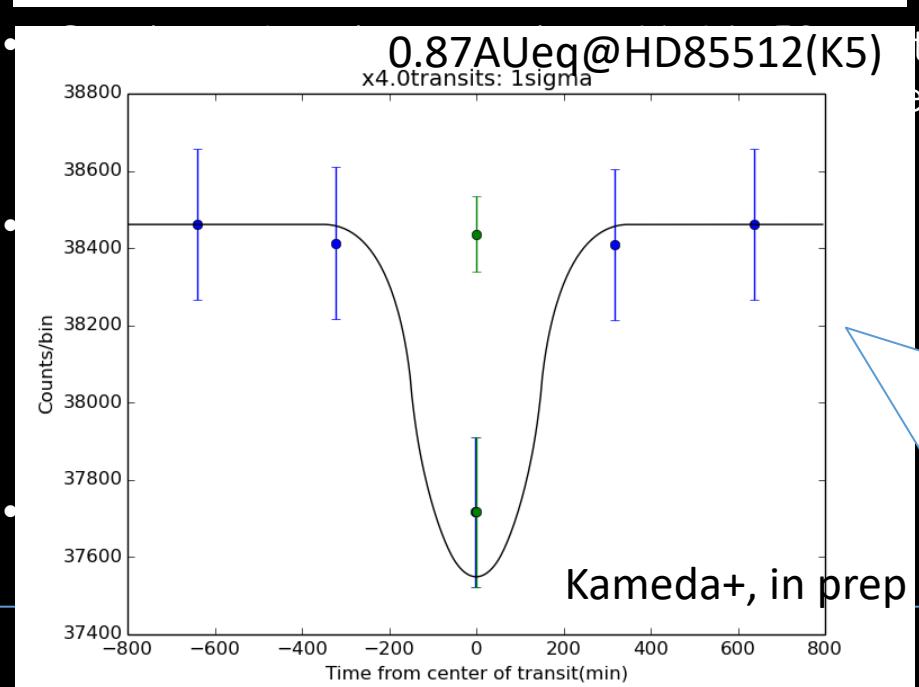
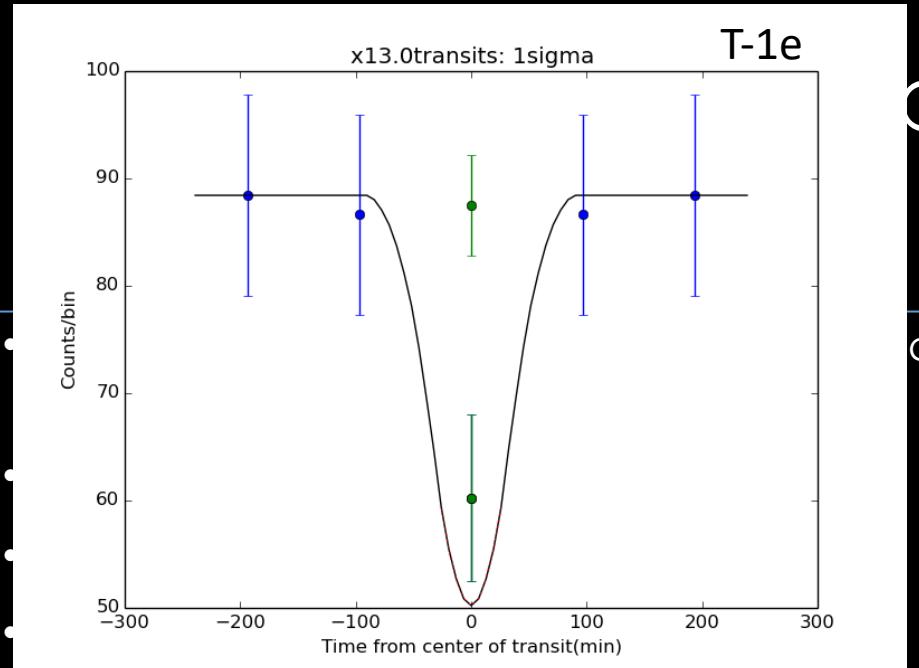
- 10-20 Earth-size exoplanets  
Oxygen <-> Ocean
- Earth-like TRAPPIST-1e can be detected in ~13 transits
- Model for upper atmosphere





UVSPEX-M





# Search for Exoplanets

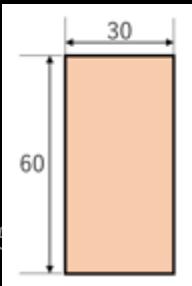
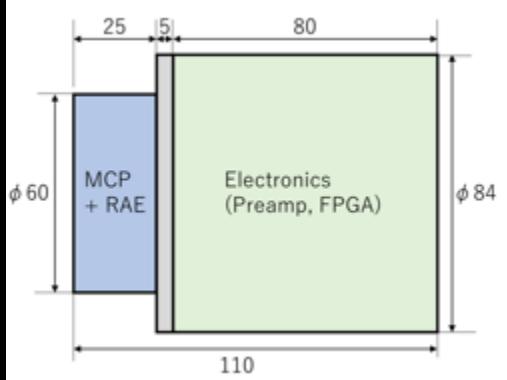
- 10-20 Earth-size exoplanets  
Oxygen  $\leftrightarrow$  Ocean
- Earth-like TRAPPIST-1e can be detected in  $\sim$ 13 transits ( $3\sigma$ )
- Model for upper atmosphere

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Earth-like planet  
 $\sim$ 4 transits ( $3\sigma$ )  
Transit depth  $\sim$ 3%  
Much higher flux

150 mm



# Telescopes for exoplanets in 2020-2030

UV

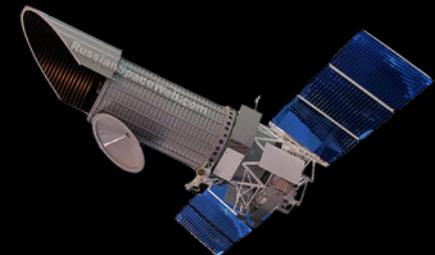
-20?? HST

2025 WSO-UV

Mid 2020 CUTE  
(cubesat, >200nm)

ESCAPE

UV irradiation, [Hydrogen, Oxygen] of terrestrial - Jupiter



## VIS/IR SPACE

2018 TESS

2019 CHEOPS

2021 JWST

2023 Twinkle

2024 Small Jasmine

2026 PLATO

2028 ARIEL

Mid 2020s WFIRST

## VIS/IR Ground

Subaru/IRD, MuSCAT

VLT/Espresso,,,

GMT, TMT, E-ELT, · · · ·

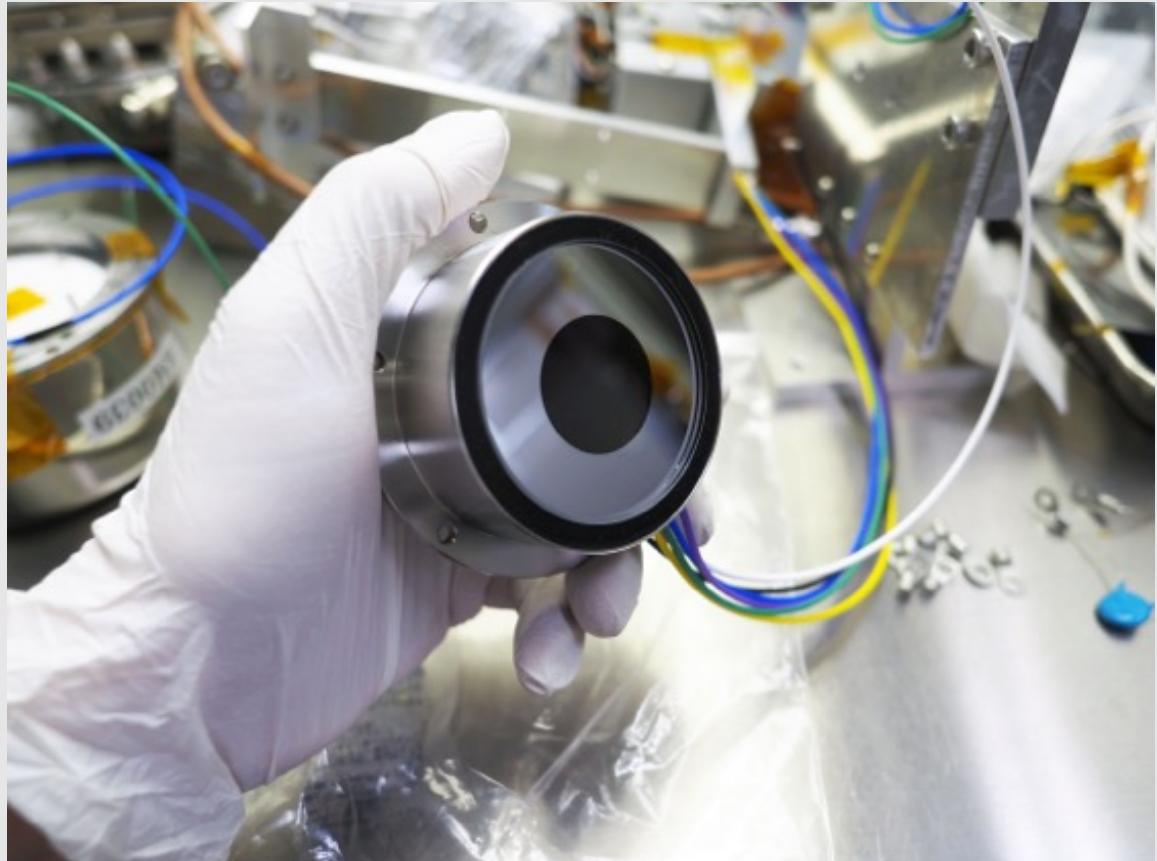
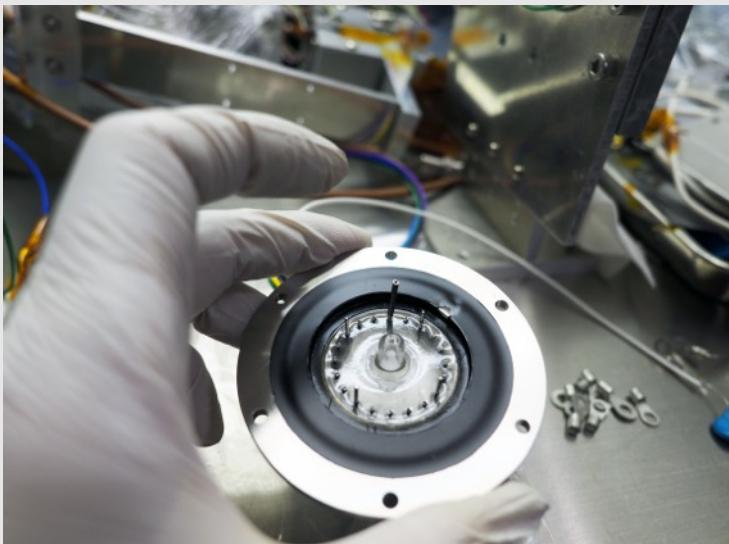


- Oxygen (+Hydrogen) Exosphere observation by VUV spectroscopy
- EUV-VUV flux is important as heat source for atmosphere
- WSO-UV is outside of oxygen geocorona
  - >Earth-like oxygen exosphere
- 2020-2030 Combination (mass, size, age, lower-upper atmosphere)

# New type MCP : Specification

## R10110M07505FR

- MgF<sub>2</sub> window : t3.0mm
- Photocathode : CsI
- MCP : 5-stage (1<sup>st</sup> :Funnel)
- Anode : RAE
- Effective Area : 16x16mm  
(smaller than previous version)



→Phosphor  
And CMOS  
readout  
For WSO-UV